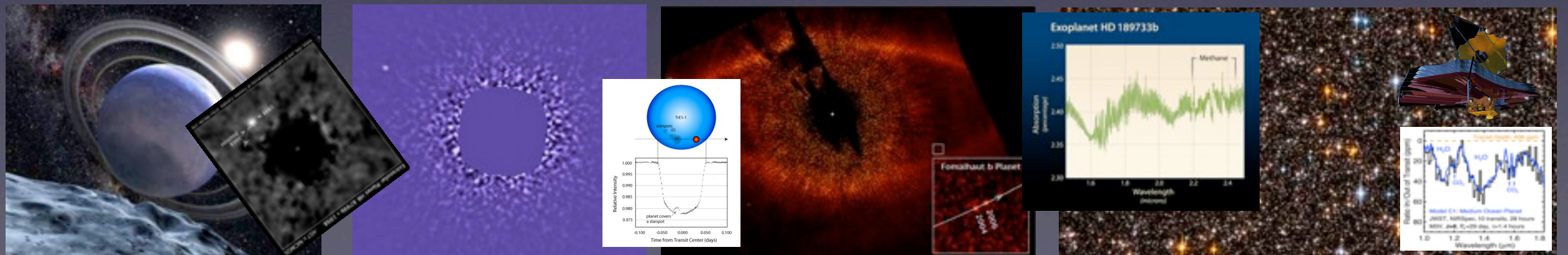


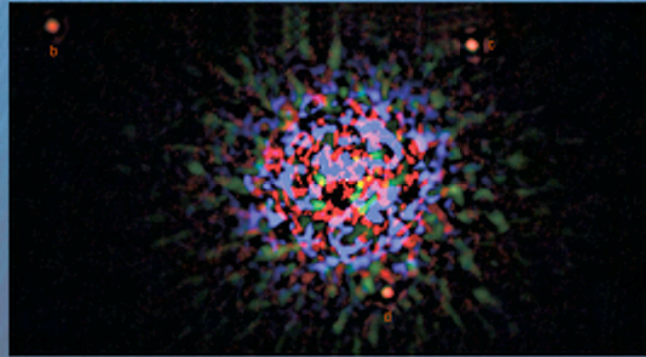
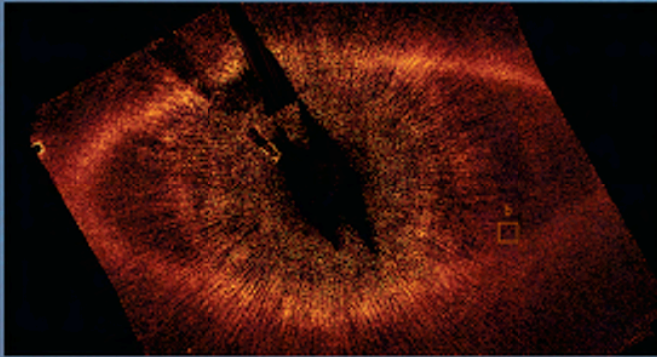
Characterizing Exoplanets - a view from Baltimore

Matt Mountain
Space Telescope Science Institute & JHU



Exoplanet Community Report

The science is **compelling**
The technology **is here**



Edited by
P.R. Lawson, W.A. Traub and S.C. Unwin

Exoplanet Community Report

The science is **compelling**
The technology **is here**

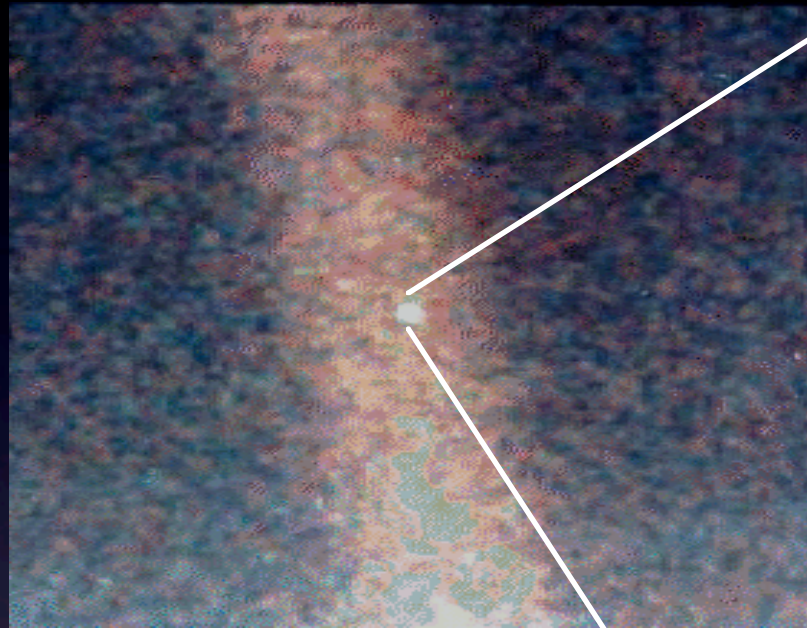
Introduction:

It is ironic that what is arguably the most compelling subject in astronomy—the search for other worlds and other life beyond our Solar System—emerges only now, in the 21st Century. Four centuries of discovery have brought us a remarkable understanding of the birth and evolution of stars, the history of galaxies, and even cosmology—the development of the entire universe, but now it seems that the first shall be last. Not for a lack of imagination or motivation, but simply for the want of technology, our oldest and deepest questions, the ones most relevant to our own origins and fate, have remained beyond our grasp for thousands of years.

We are indeed fortunate to live in the time when this last barrier to our search is falling. It

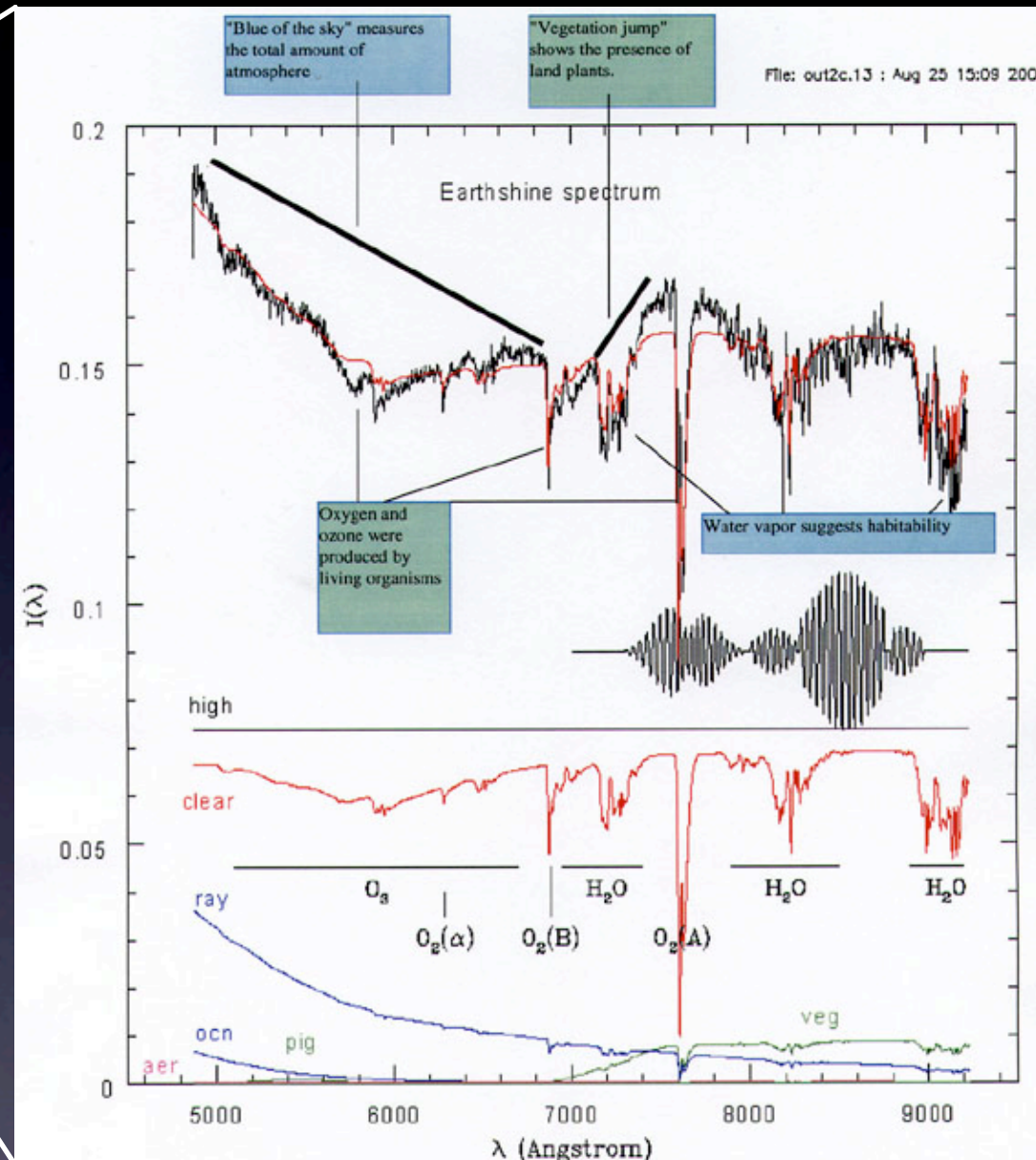
is the technology here ?

assumption 1 of 2



Is there life on this planet?

We require spectra
of 29 ~30 m_v sources



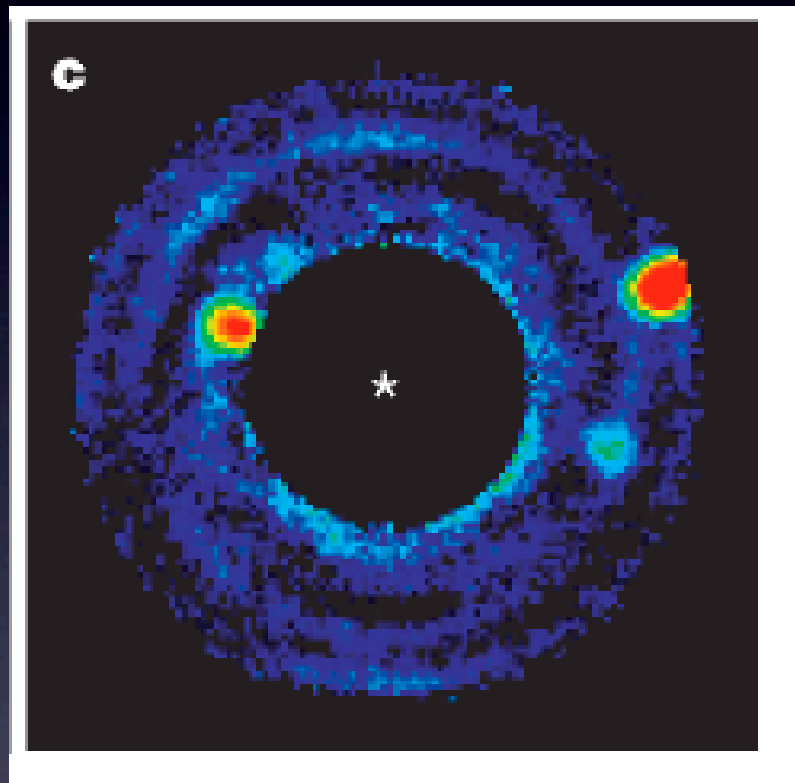
By 2020 “pale blue dots” will be necessary but not
sufficient scientific motivation **we will need spectra**

is the technology **here** ?

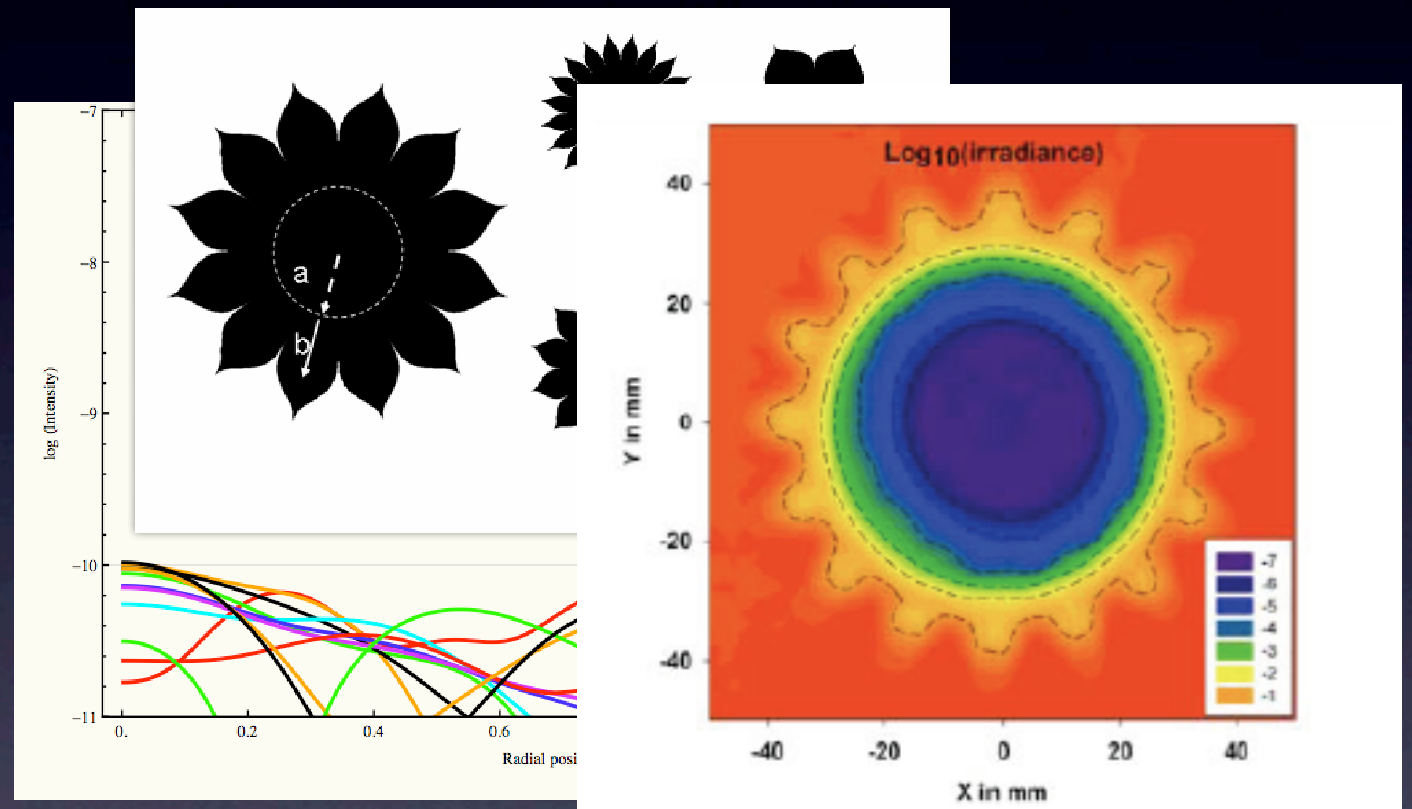
assumption 2 of 2

Coronagraph

Free flying star-shade



Trauger & Traub Nature Letters, April 2007



Cash, Soummer, various, 2008, Leviton et al. 2007; Schindhelm et al. 2007

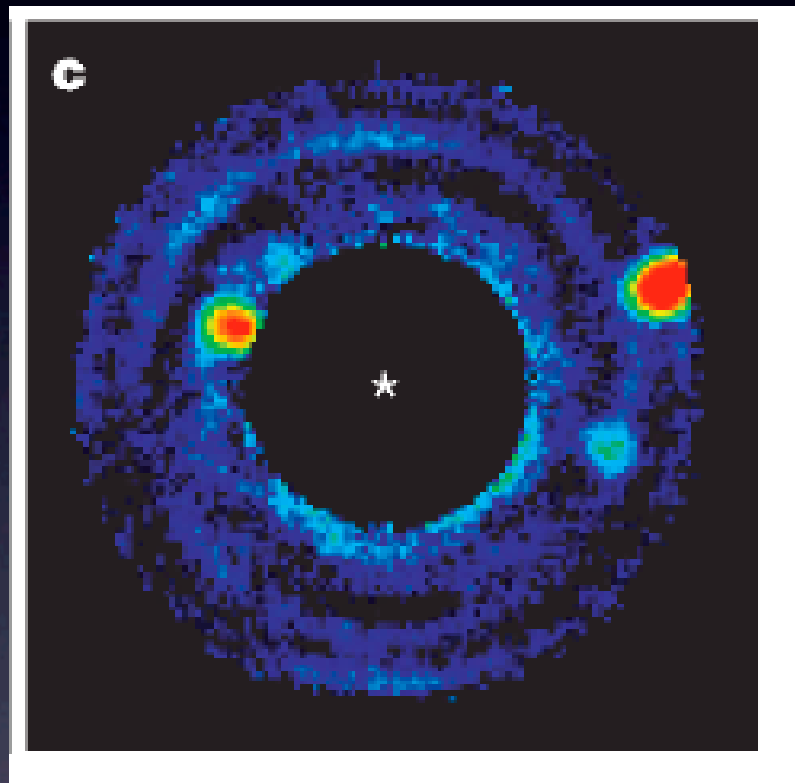
Starlight suppression to $\sim 10^{-10}$
is a **solvable problem from Space**

is the technology **here** ?

assumption 2 of 2

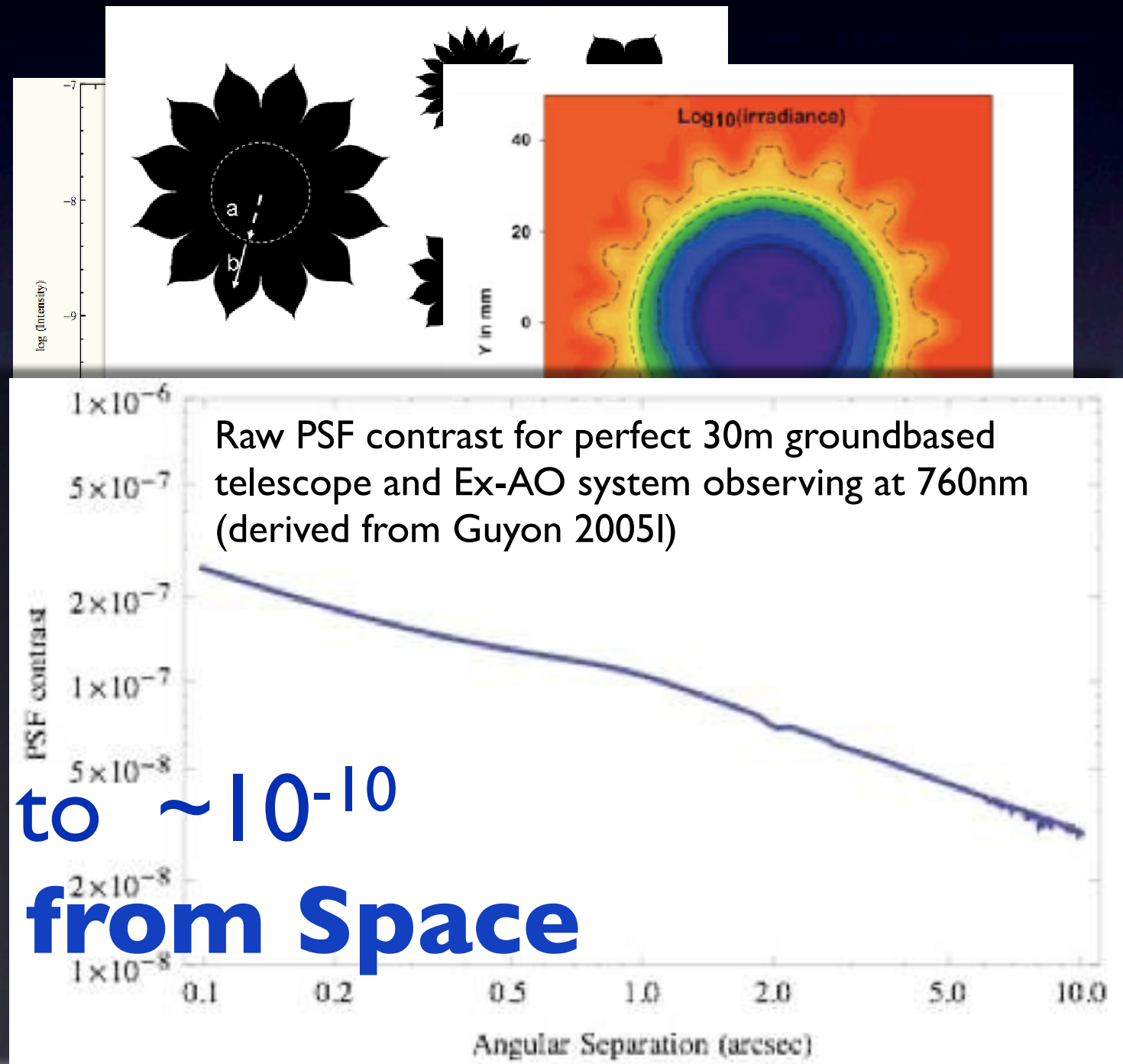
Coronagraph

Free flying star-shade



Trauger & Traub Nature Letters, April 2007

Starlight suppression is a **solvable problem** to $\sim 10^{-10}$ **from Space**



Observable Drake Equation – after Reid & Hawley

$N_{L,T}$ is the number of life bearing planets at time T

$$N_{L,T} = N_{*,T} p_p n_e p_w p_l$$

Observable

Number of stars @ T:

$$N_{*,T} = SRF(t) * \Psi(m) * \Lambda(m,t)$$

Prob. of planet system:

$$p_p = f(Z, m)$$

No. of terrestrial planets

$$n_e = n(0.1 m_e < m_p < 10 m_e)$$

Prob. of liquid water

$$p_w = f(\bar{n}_b, \epsilon, r_{orbit}, L^*)$$

Prob. of life

p_l

Observable Drake Equation – after Reid & Hawley

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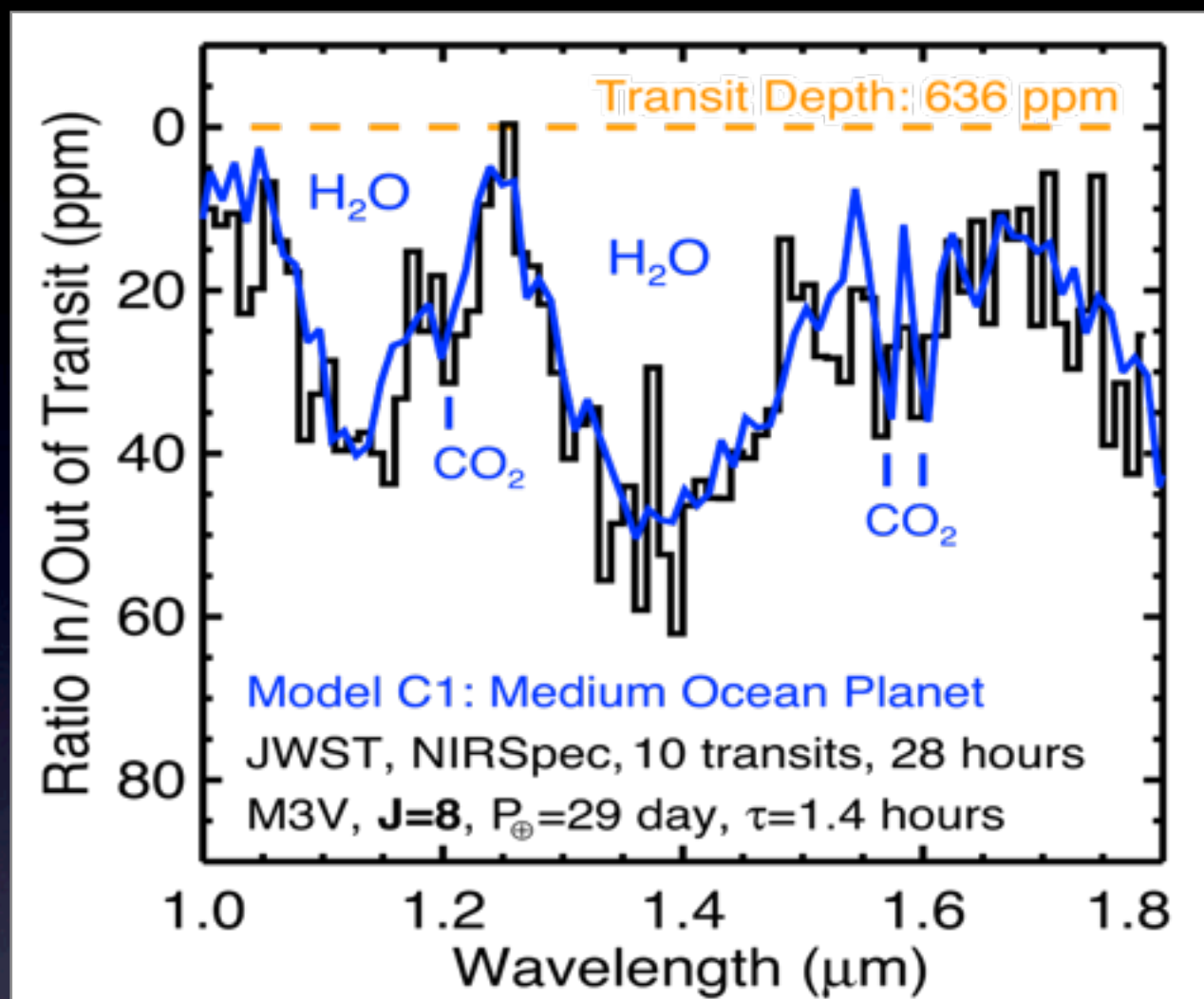


Prob. of life

p_l



Transit Spectra of a Habitable Ocean Planet



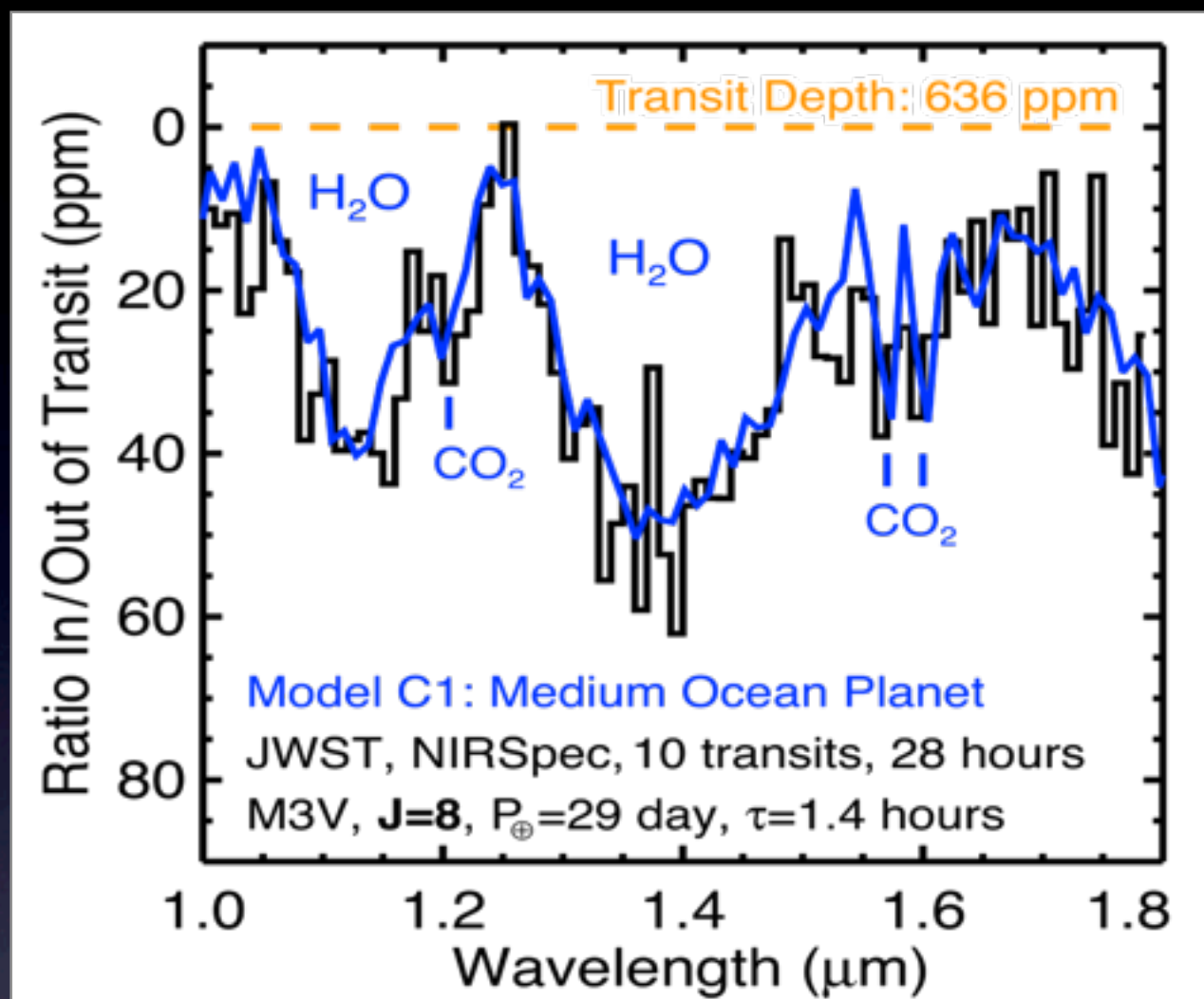
Gliese 581 (M3V, J=6.7)

- b: 5.4 days, 15.6 M_⊕
- c: 12.9 days, 5.1 M_⊕
- d: 83.4 days, 8.3 M_⊕

Find one that transits...

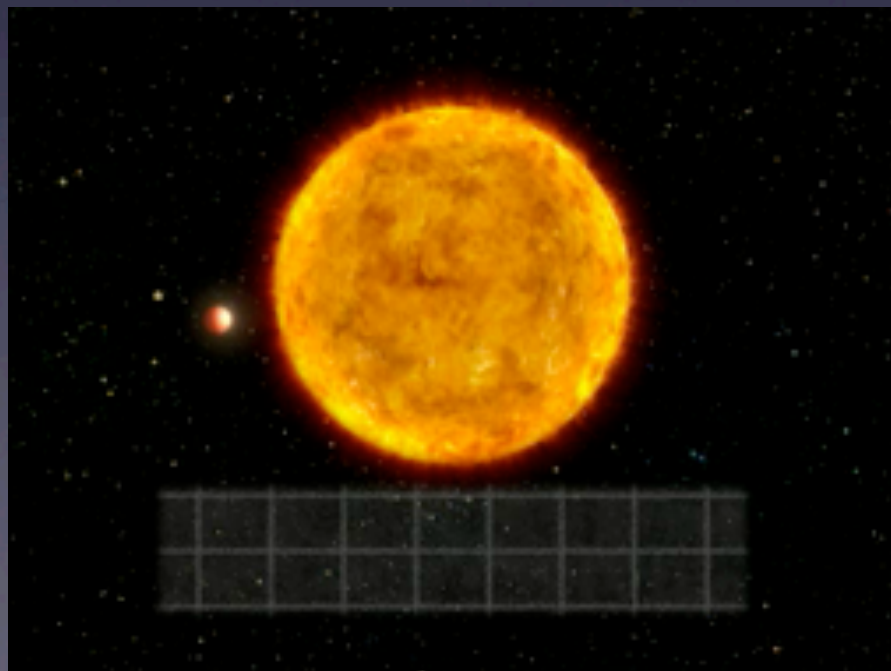
- 6000 M dwarfs with J<10
- Habitable → 11% transit
- Up to 70 transits for J<10

Transit Spectra of a Habitable Ocean Planet



Gliese 581 (M3V, J=6.7)

- b: 5.4 days, $15.6 M_{\oplus}$
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Find one that transits...

- 6000 M dwarfs with $J < 10$
- Habitable \rightarrow 11% transit
- Up to 70 transits for $J < 10$

how big a space
telescope do we need to
solve the **Observable
Drake Equation?**



$$N_{L,T} = N_{*,T} p_p n_e p_w p_I$$

Number of stars @ T:

Prob. of planet system:

No. of terrestrial planets

Prob. of liquid water

$$N_{*,T} = SRF(t) * \Psi(m) * \Lambda(m,t)$$

$$p_p = f(Z, m)$$

$$n_e = n (0.1 m_e < m_p < 10 m_e)$$

$$p_w = f(\bar{n}_b, \varepsilon, r_{orbit}, L_*)$$

Observable



$$N_{L,T} = N_{*,Observed} \cdot \eta_{earth} \cdot p_I$$

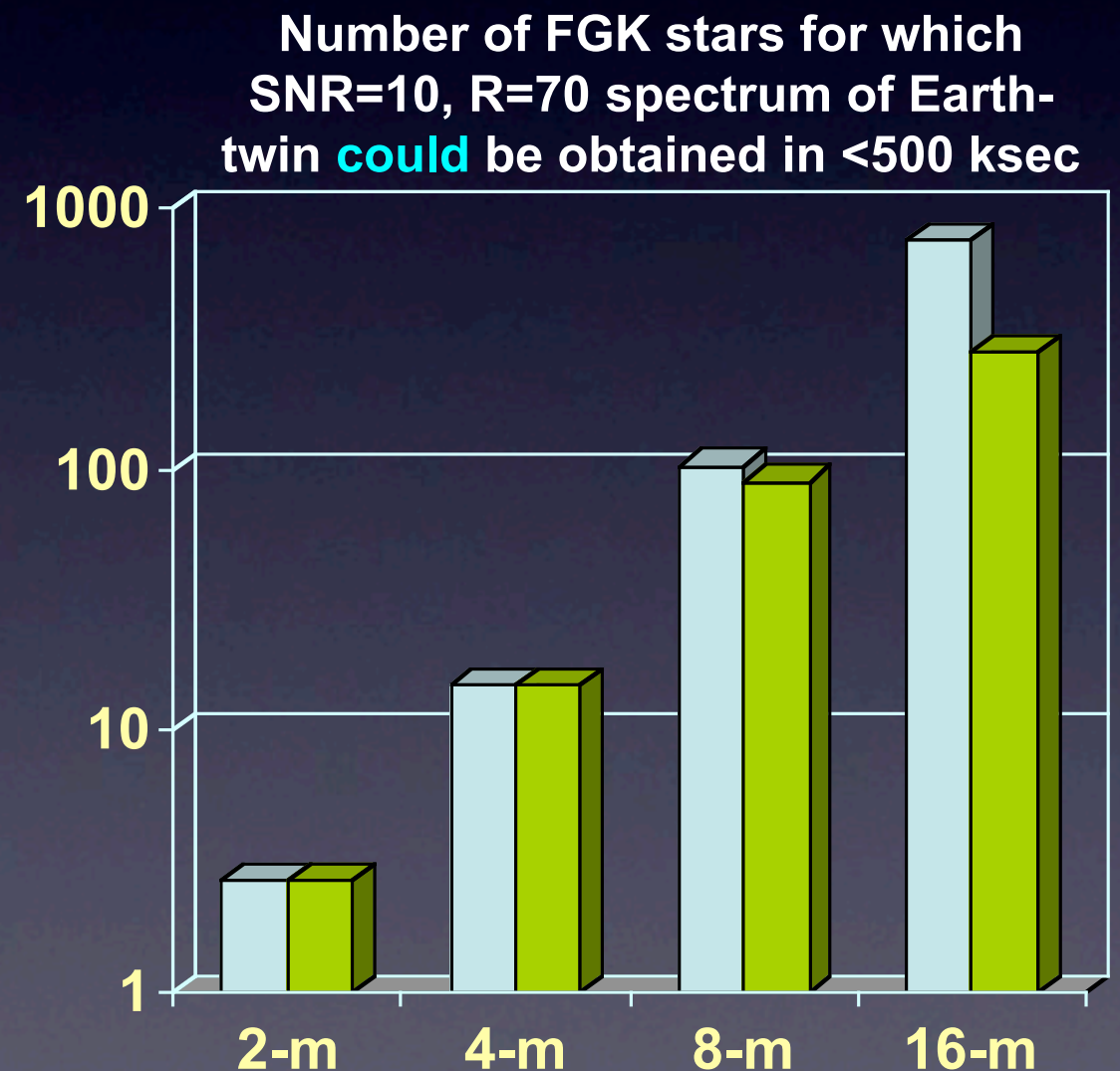
We know where all the stars are

- how many could we search for earths ?

Green bars show the number of FGK stars that could be observed 3x each in a 5-year mission without exceeding 20% of total observing time available to community.

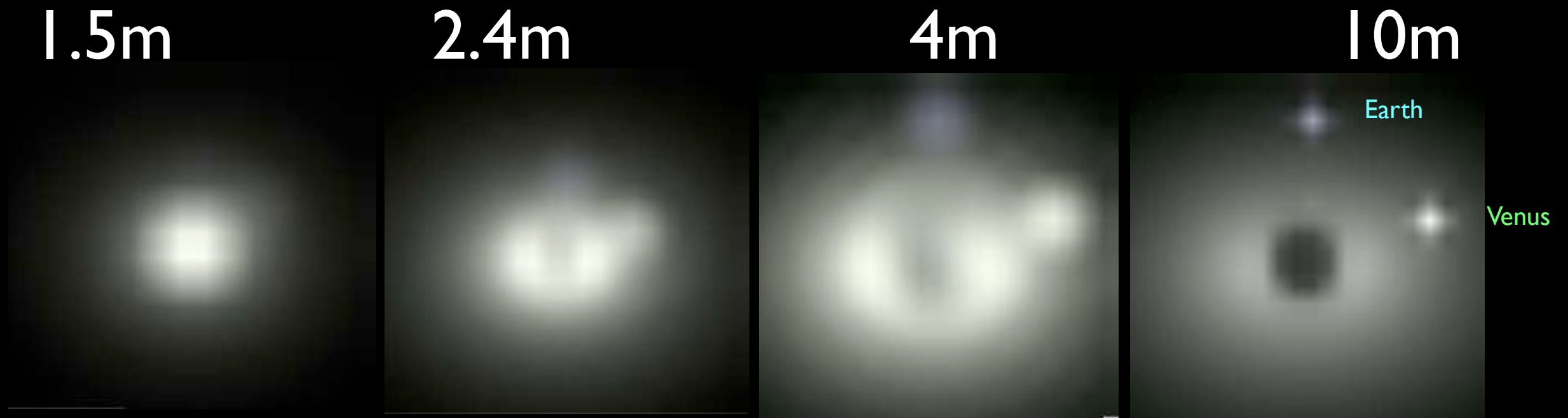
Need to multiply these values by $\eta_{\text{Earth}} \times f_B$ to get the number of potentially life-bearing planets detected by a space telescope.

η_{Earth} = fraction of stars with Earth-mass planets in HZ
 f_B = fraction of the Earth-mass planets that have detectable biosignatures



Characterizing Exoplanets from Space

Credit: Web Cash 2008



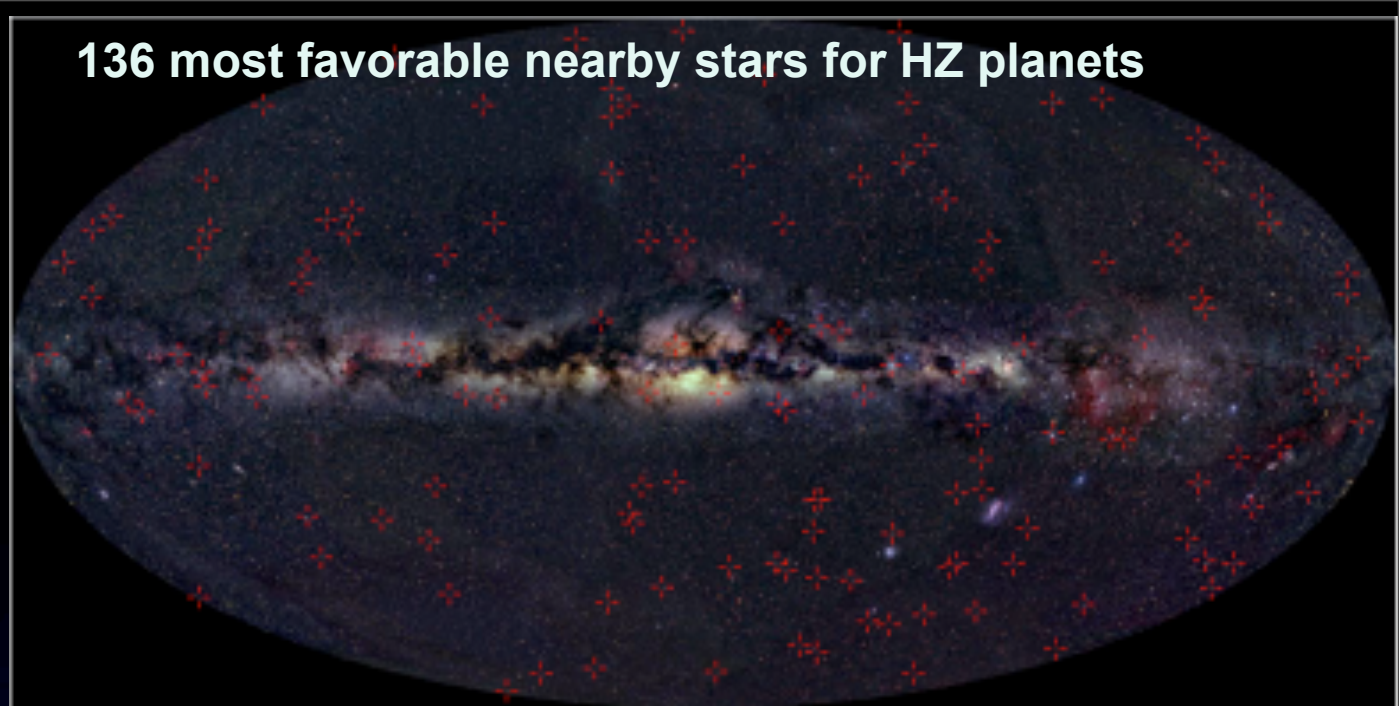
Above: a simulation of our solar system at a distance of 10 pc observed with an external occulter and a telescope with the indicated aperture size. The two planets are Earth and Venus.

characterizing, and discriminating terrestrial scale planets
from their parent star
requires aperture **and** angular resolution

Courtesy ATLAST Team

how big a space
telescope do we need
to solve the **Observable
Drake Equation?**

136 most favorable nearby stars for HZ planets



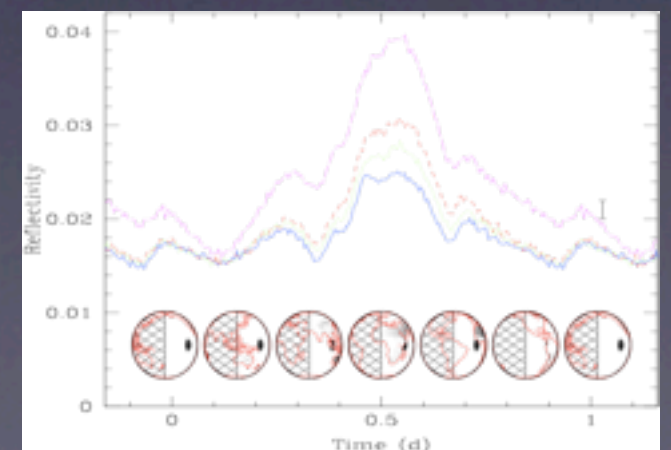
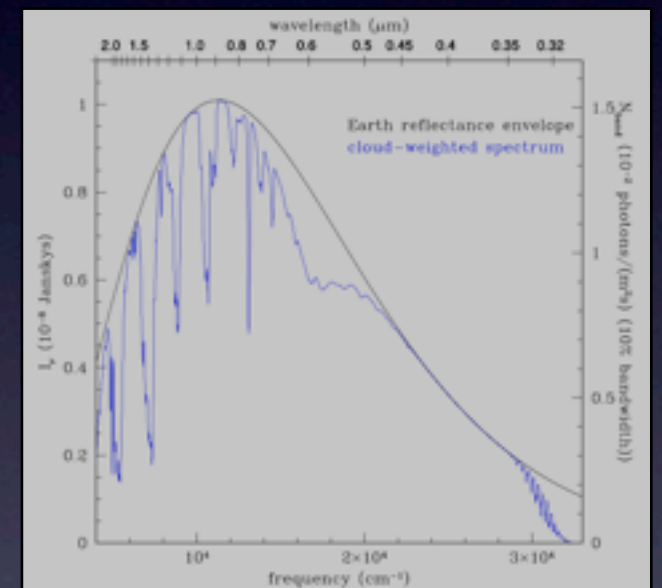
$$N_{L,t} \sim D_{Tel}^3 \cdot \eta_{earth} \cdot p_L$$

If: $\eta_{earth} \cdot p_L \sim 1$ then $D_{tel} \sim 4m$

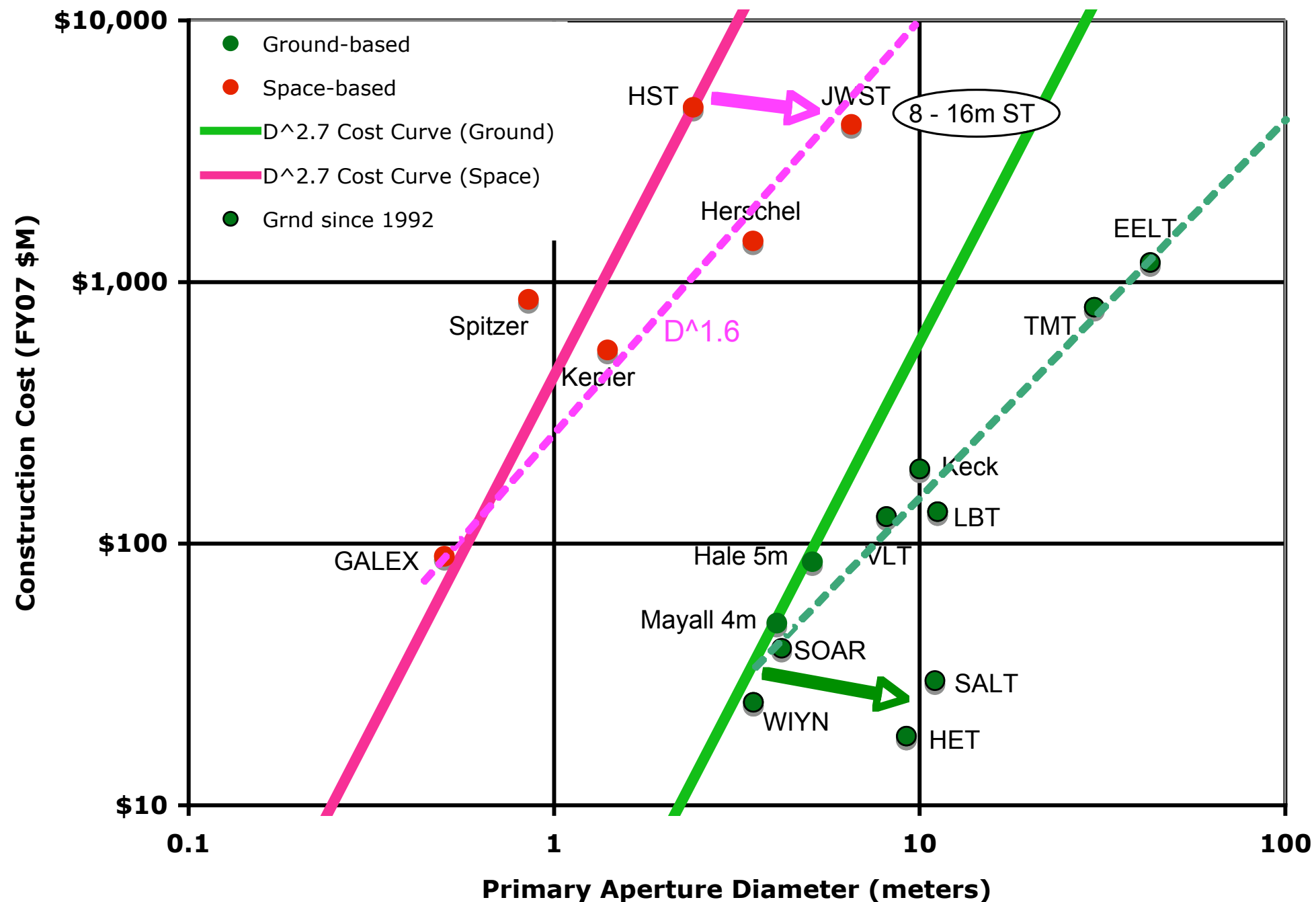
$\eta_{earth} \cdot p_L < 1$ then $D_{tel} \sim 8m$

$\eta_{earth} \cdot p_L \ll 1$ then $D_{tel} \sim 16m$

(assuming fixed zodiacal contribution)



cost does **not** follow a fixed scaling relation with aperture as technology or architecture advance



for example

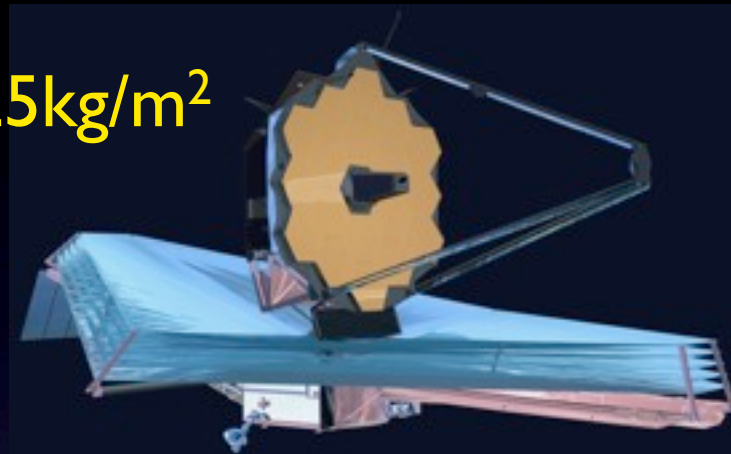
HST 2.4m



$\sim 140 \text{ kg/m}^2$

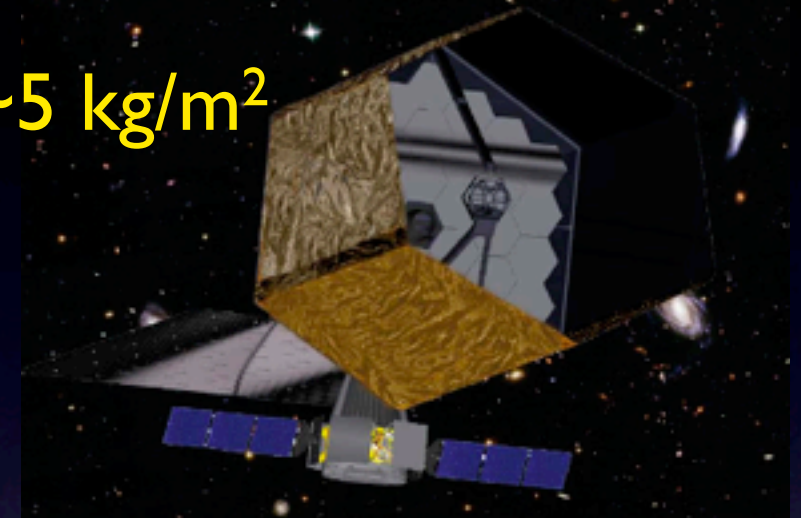
JWST 6.5m

$\sim 25 \text{ kg/m}^2$



8m~16m LST

$\sim 5 \text{ kg/m}^2$



Passive
control



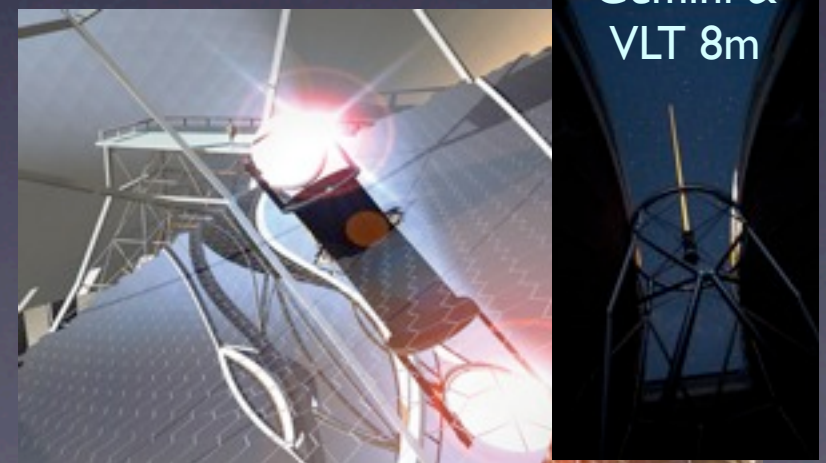
Palomar 5m

Active
control



Keck 10m

Fully active and
adaptive control



E-ELT/TMT 30m~40m

Gemini &
VLT 8m

pathways to a large UV-IR space telescope

If Ares V is built by 2019 ...



8-m monolithic
mirror Telescope
in ~2025

and/or

16-m segmented
mirror Telescope
in ~2030+

If Ares V is not built ...



Delta IV HLV



9.2-m segmented
mirror Telescope
in ~2028

or

Elliptical (light-
weight) monolithic
mirror Telescope
in ~2028

Ares V payload to L2 = 65 mT, Delta IV HLV payload to L2 = 16 mT

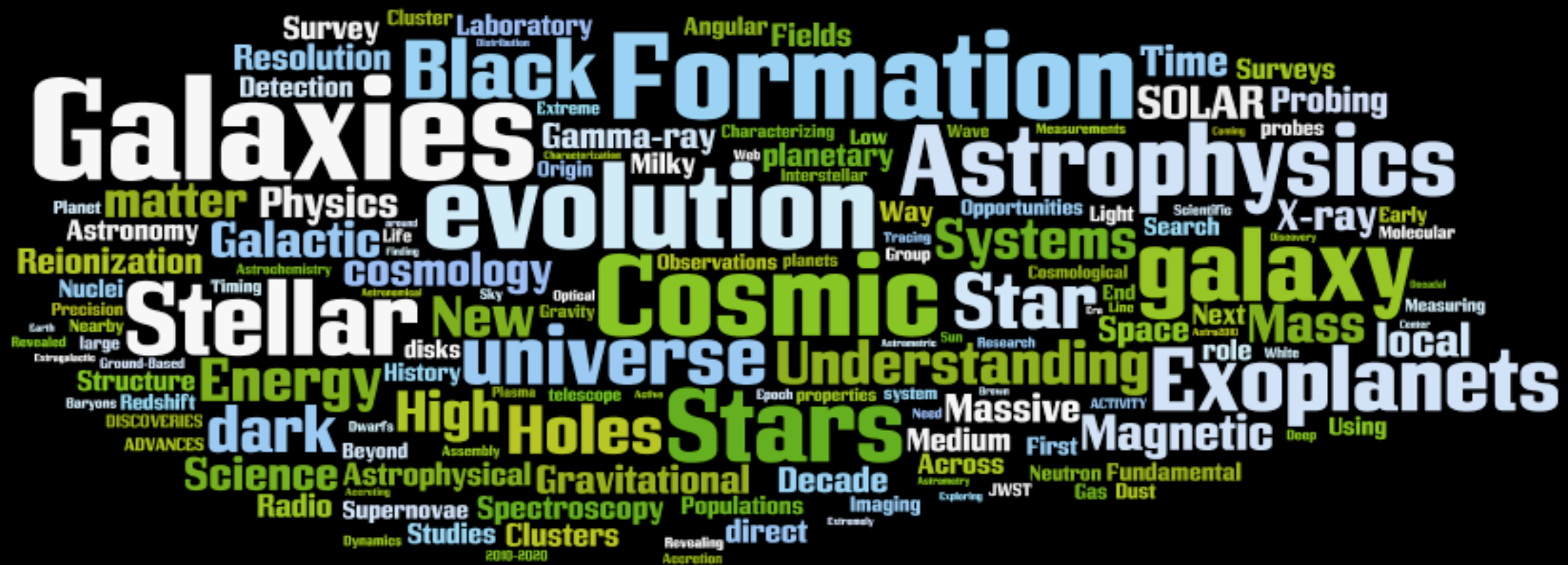
The science is **compelling**

“Four hundred years after Galileo’s discoveries via the first astronomical use of the telescope, the world’s astronomers are once again using powerful new instrumentation to make startling discoveries of and about new planets, quite literally other worlds, **which promise to once again transform our understanding of the nature of the Earth and of life’s and humanity’s places in the cosmos.**”

Ed Turner, 24th March 2009

The Search for Life in the Universe @ STScI

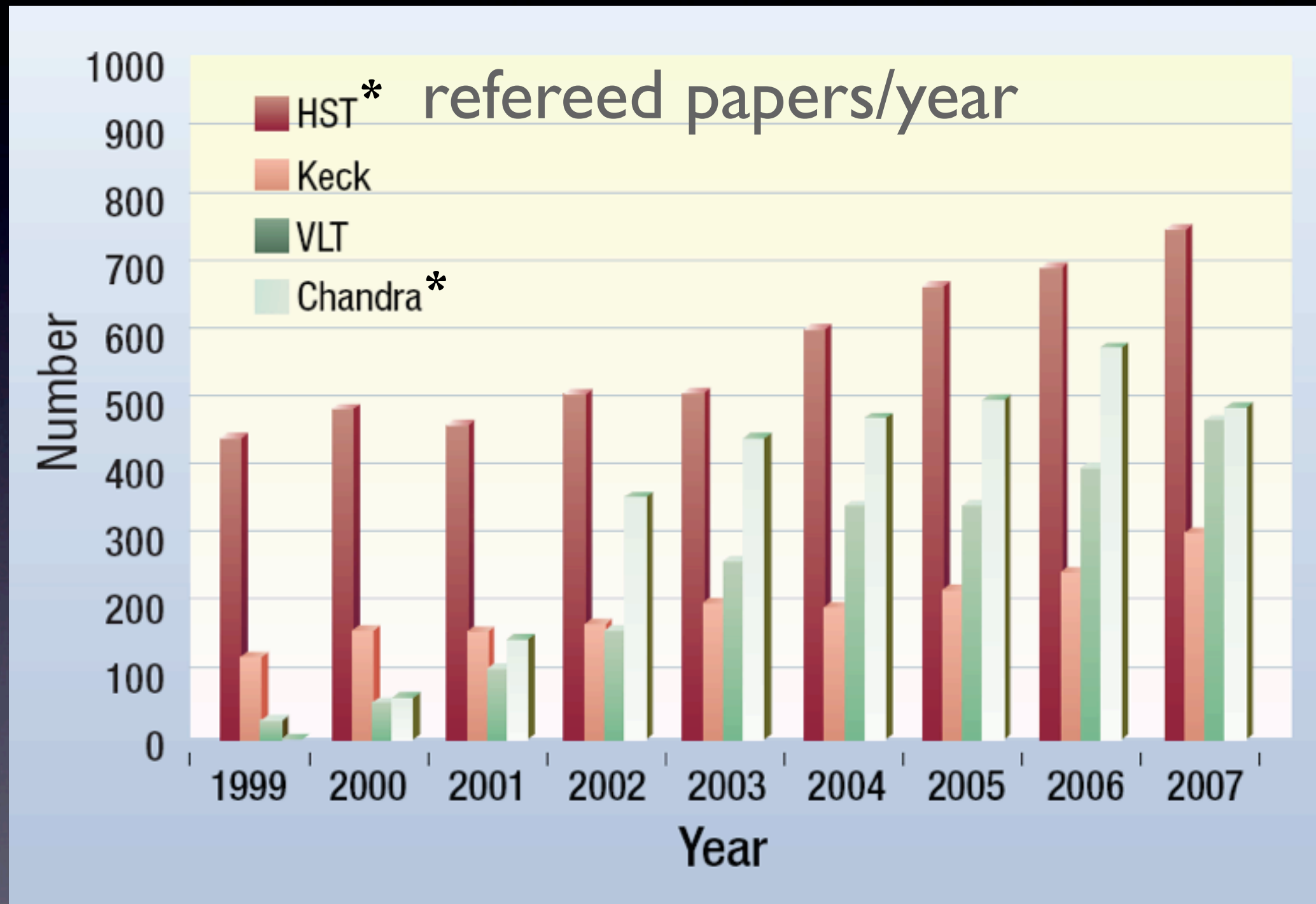
however... summary of 132 science
white papers submitted to Astro2010



Courtesy of <http://www.wordle.net/>, IBM Research media communications lab. - uniform use of “Exo-planets”

~85% of submissions **do not** reference “exo-planets”

the most productive “Flagships” have broad community participation **hence support**



* for these mission NASA provides substantial grant support to the community

observational astrophysics need flagships

Signal
Noise

\propto Telescope Diameter
Image size

$$\times \sqrt{\frac{QE_{\lambda}}{B_{\lambda}}}$$

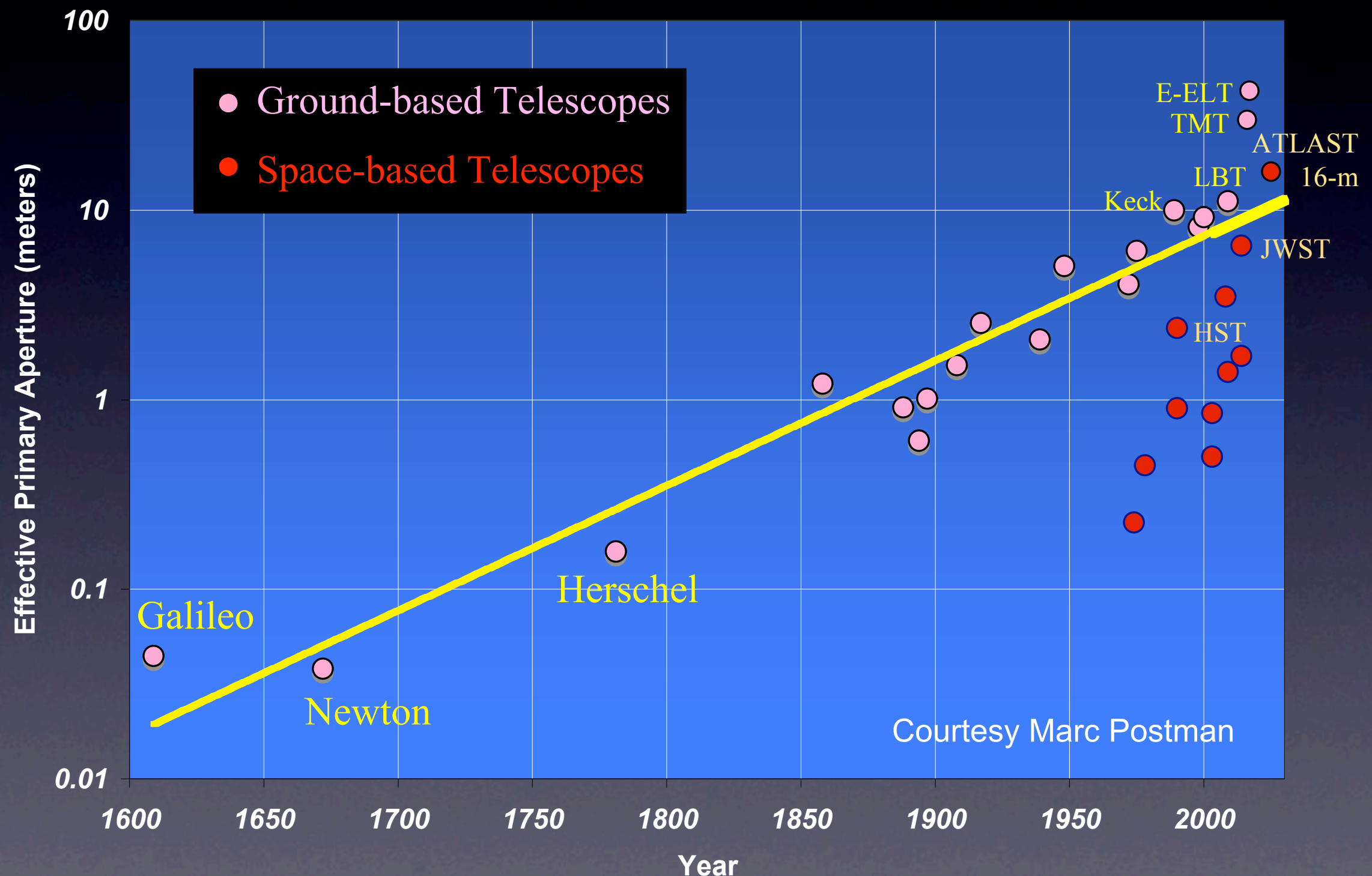
QE = detector quantum efficiency $\sim 100\%$

B = sky, telescope background \sim low in space

Image size, over large FOV \sim small in space

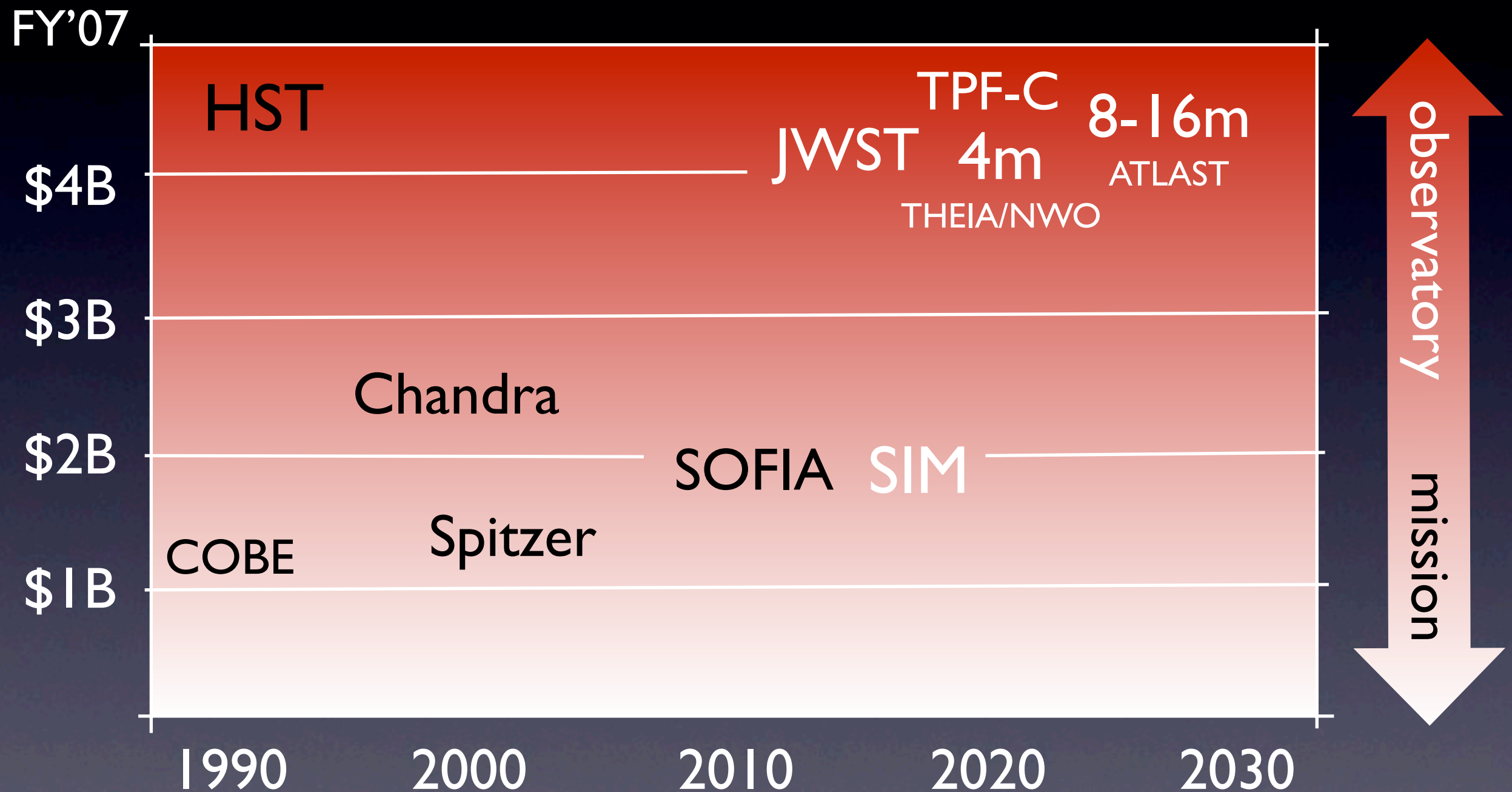
Observational astrophysics remains a photon limited science and hence there is a compelling case for large aperture UV, Optical & Near-infrared telescopes **in space**

aperture growth driven by science, technological maturity
but also by philanthropy, science and/or industrial policy...

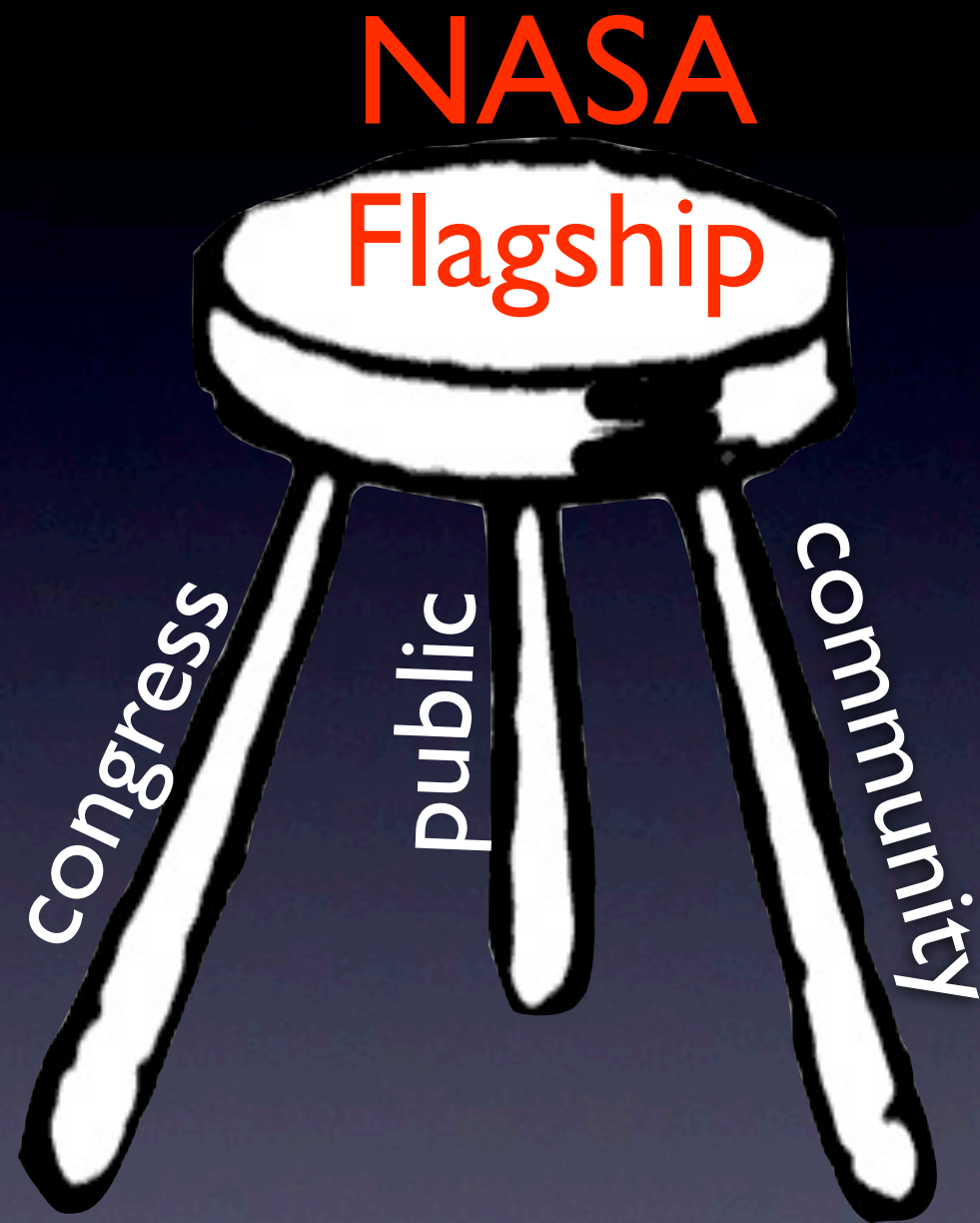


Primary Telescope Aperture vs. Time

flagship cost and 'expectations'



huge competition for the few slots in the **top-right corner**



NASA

Flagship

congress

public

community

4680

1965

601

525

19

117

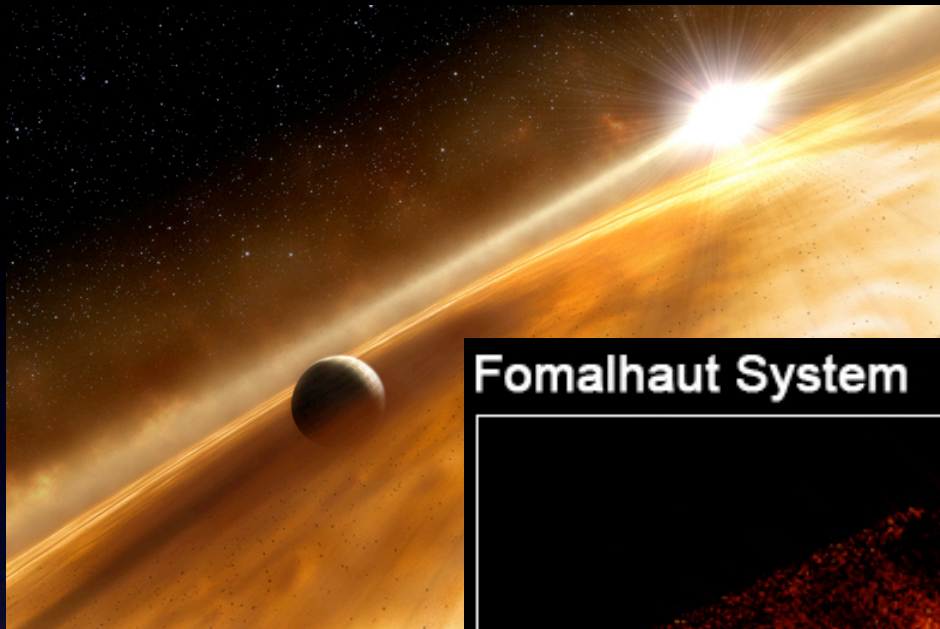
HST Archive Users 2007

June 15, 2020



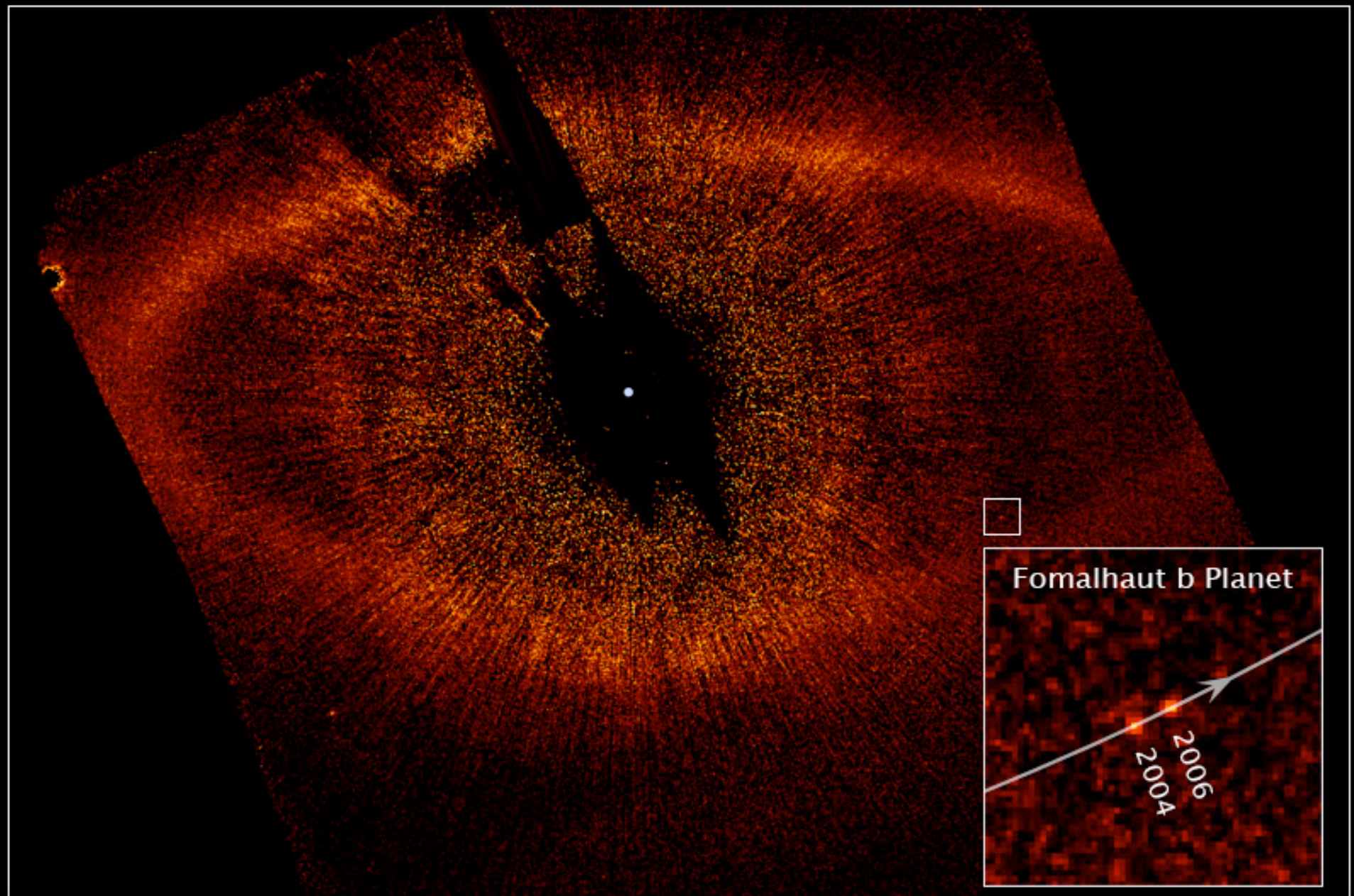
www.time.com

Hubble images its first planet outside our solar system



Fomalhaut System

Hubble Space Telescope • ACS/HRC

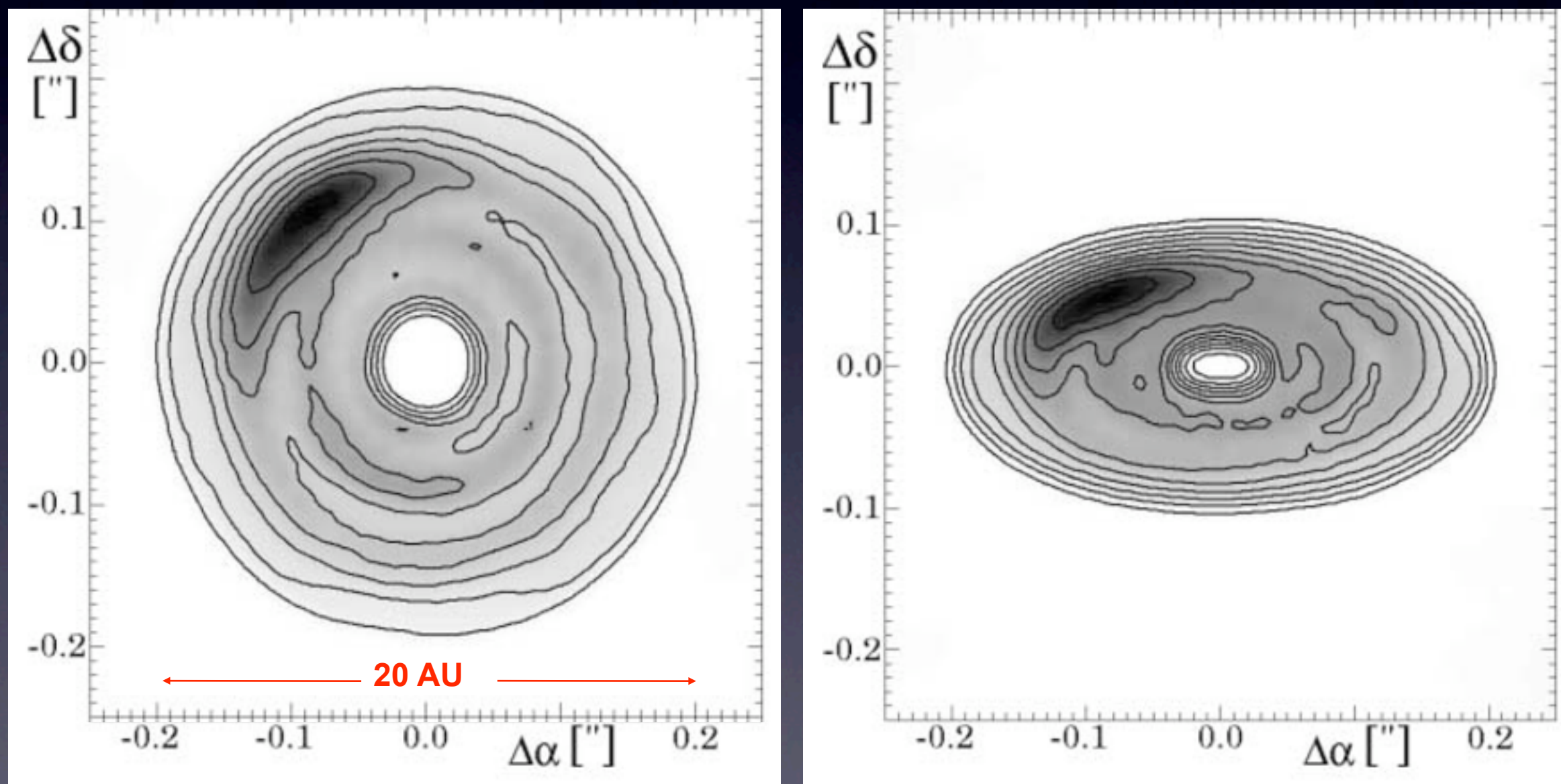


NASA, ESA, and P. Kalas (University of California, Berkeley)

STScI-PRC08-39a

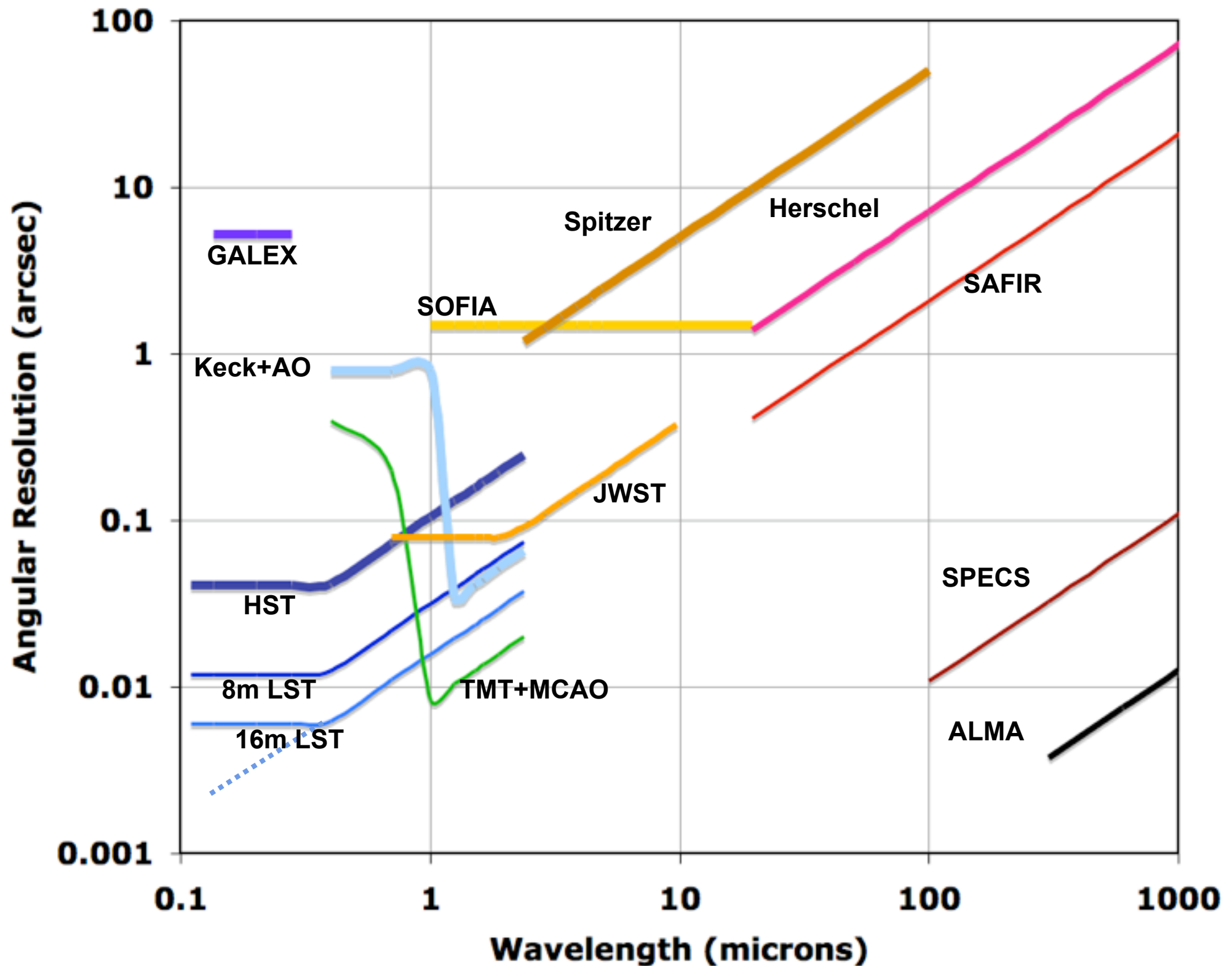
example: early evolution of proto-planetary disks

Simulated ALMA 900 GHz images of protoplanetary disk at distance of 50 pc (Wolf & Klahr ApJ, 2002)



From Marc Postman

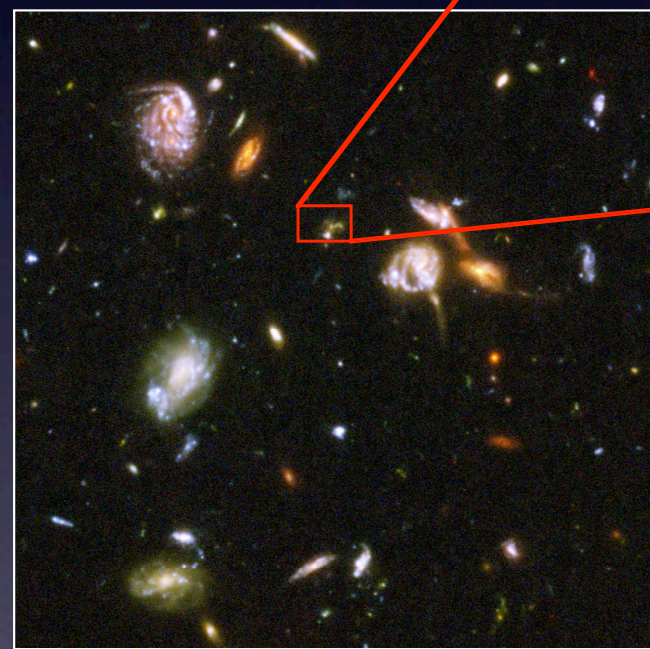
Multi-wavelength Angular Resolution



From Marc Postman

Figure derived from ExoPTF (Lunine et al. 2008)

“modern” galaxy evolution



HST Ultra Deep Field

Faint Galaxy:

25.1 AB mag (330 nJy) in I-band
0.75 arc seconds across
2 “peaks” in light distribution
Morphology unknown

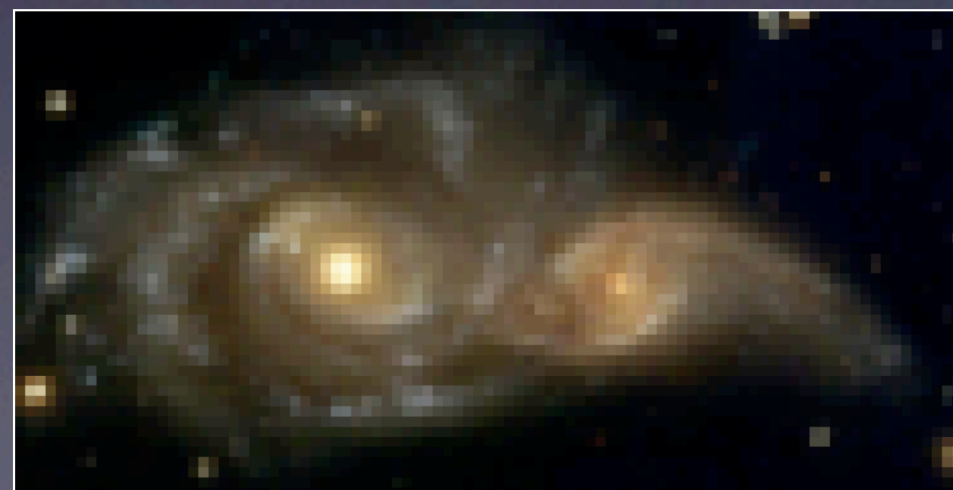
Courtesy ATLAST Team



HST 2.4-m,
t~900 ksec



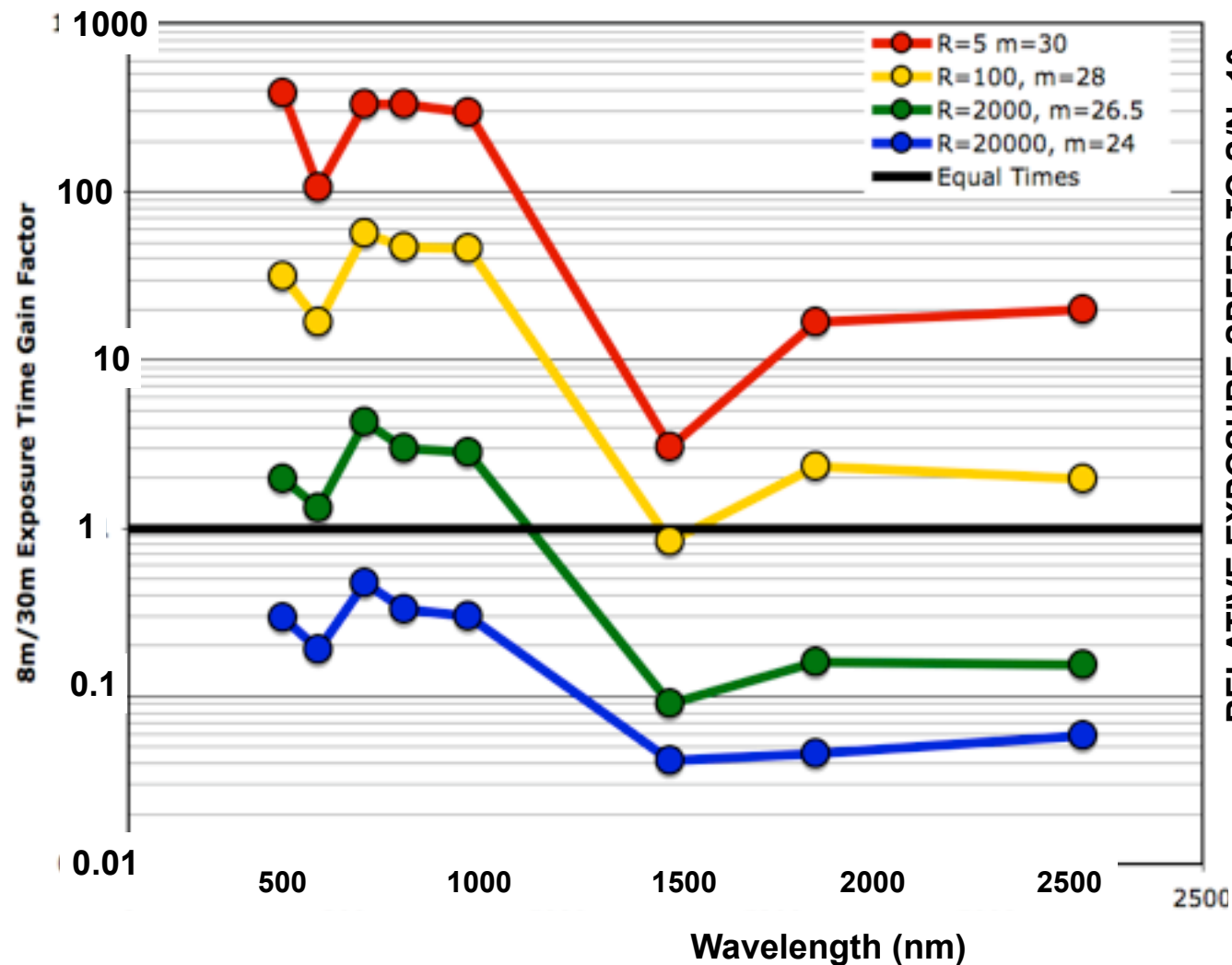
8-m LST,
t~25 ksec



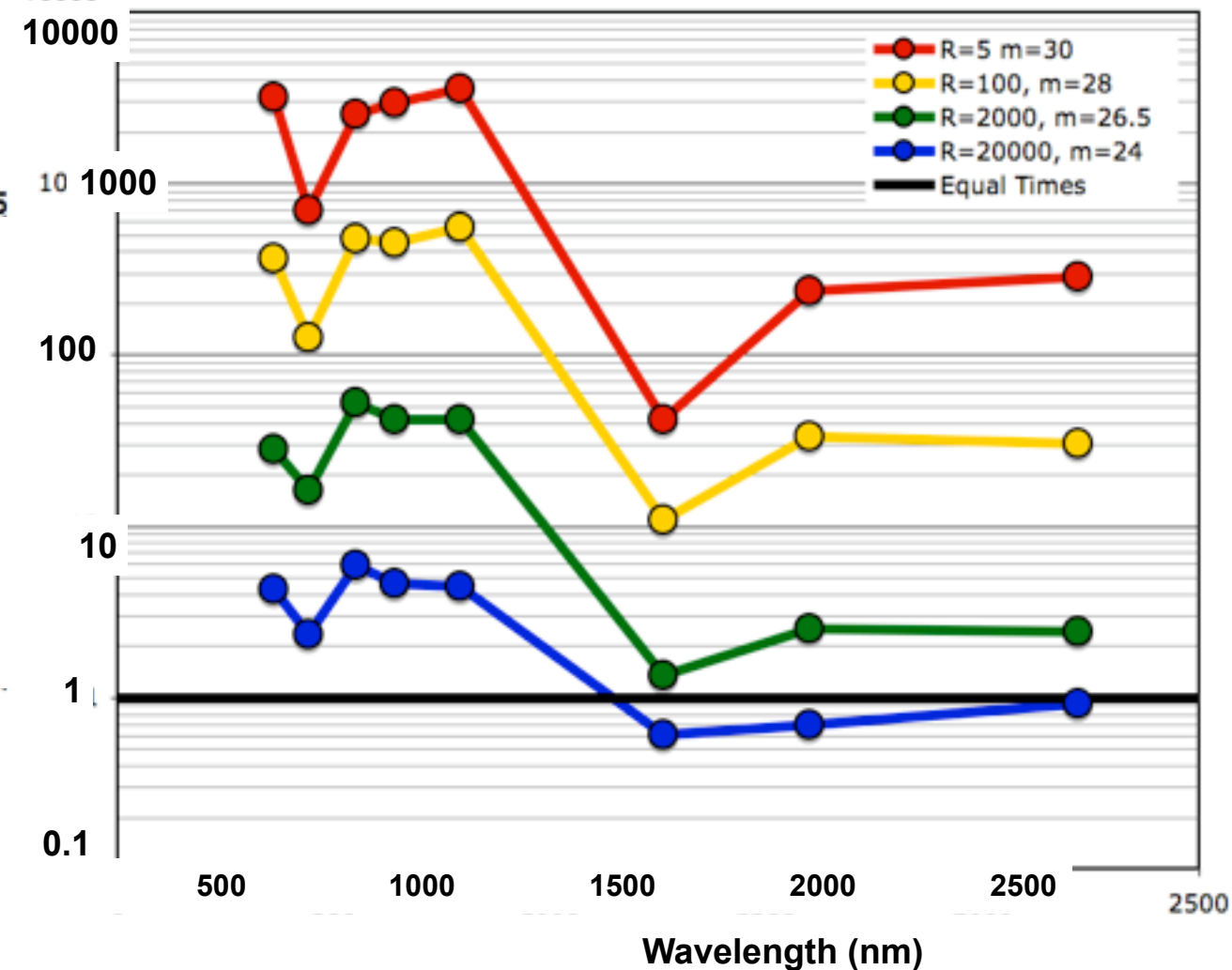
16-m LST,
t~3 ksec

time gain factor

8-m ST vs. 30-m Ground-based+AO in NIR

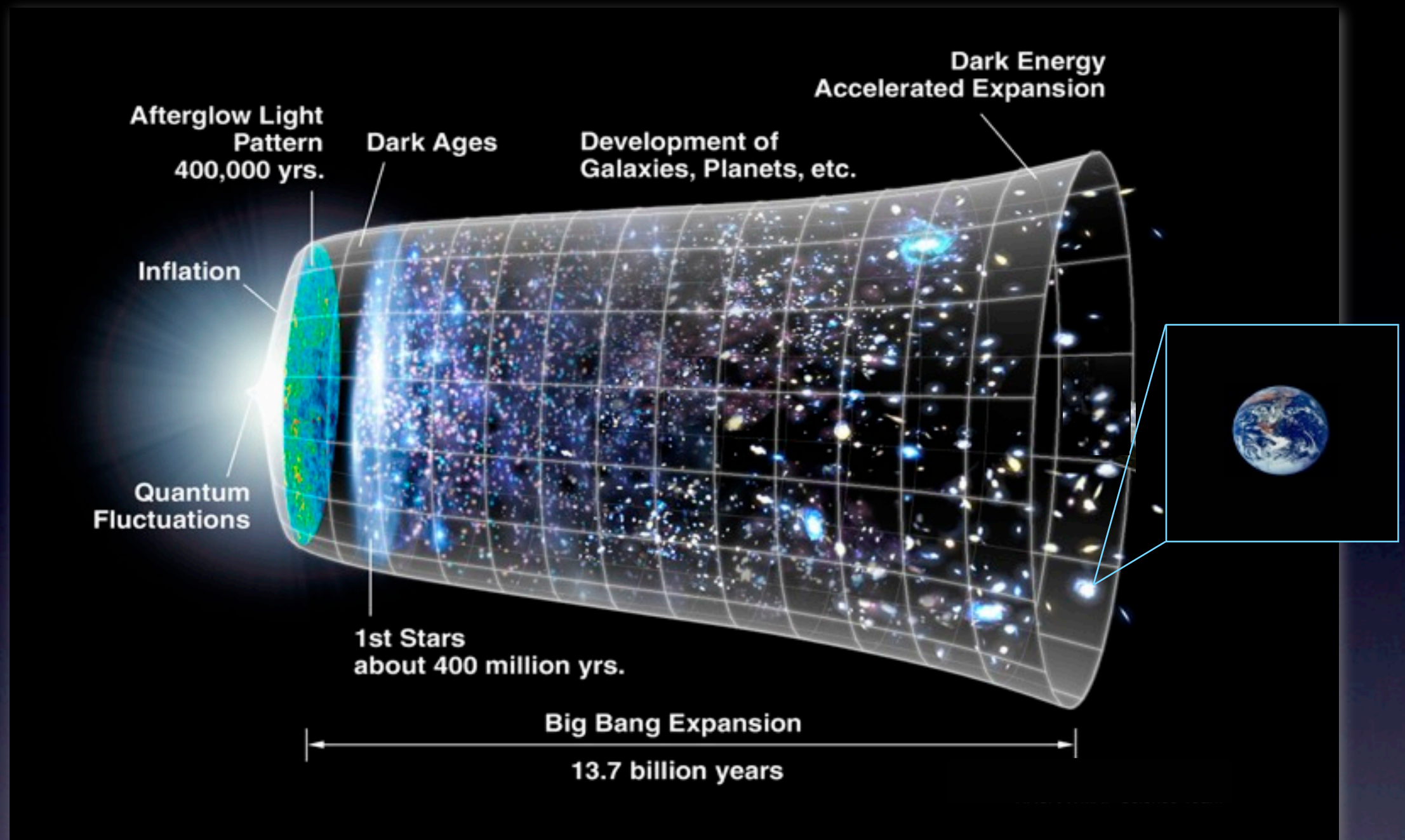


16-m ST vs. 30-m Ground-based+AO in NIR



8-m ST faster than 30-m on ground for all imaging and for most $R=100$ low-res spectroscopy. 8-m also faster for med-resolution spectroscopy in optical band.

16-m ST faster than 30-m on ground for all imaging and spectroscopy except when $R > \text{few} \times 1000$ in the NIR. Unique parameter space for hi-res spectroscopy in optical band.

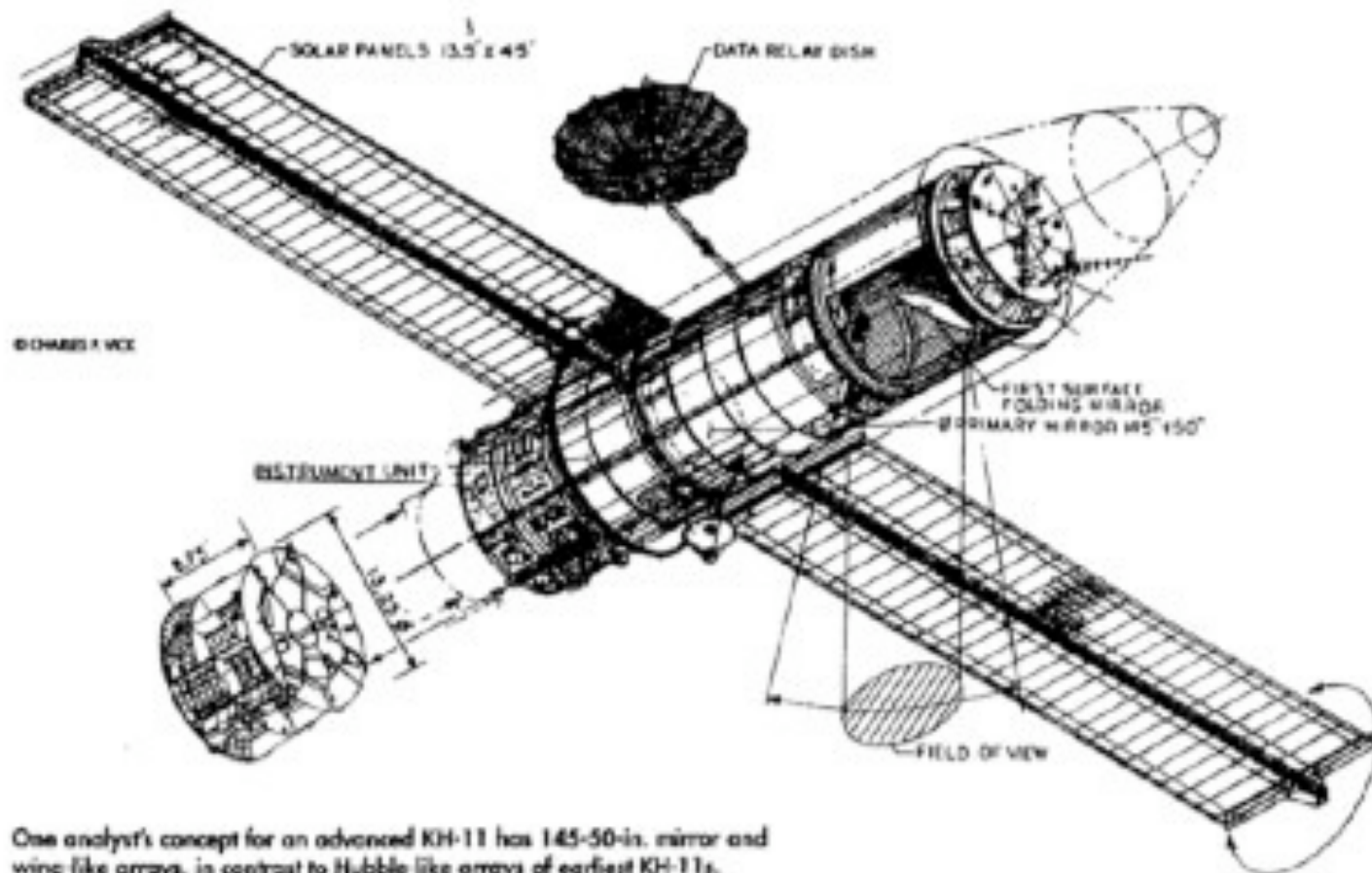


the most compelling science themes, the search for life
and
unraveling the complexity of the modern Universe - **which led**
from the Big Bang to the emergence of a sustainable, habitable planet,
require a generation of new and powerful space telescopes

Space Science has always built on investments made “elsewhere”

HEADLINE NEWS

Advanced KH-11 Broadens U.S. Recon Capability



One analyst's concept for an advanced KH-11 has 145-50-in. mirror and wing-like arrays, in contrast to Hubble-like arrays of earliest KH-11s.

24 AVIATION WEEK & SPACE TECHNOLOGY/JANUARY 6, 1997

The two main contractors that built the telescope had allegedly extensive experience building this kind of spacecraft - but not much is known publicly about these programs.

Space Science has always built on investments made “elsewhere”

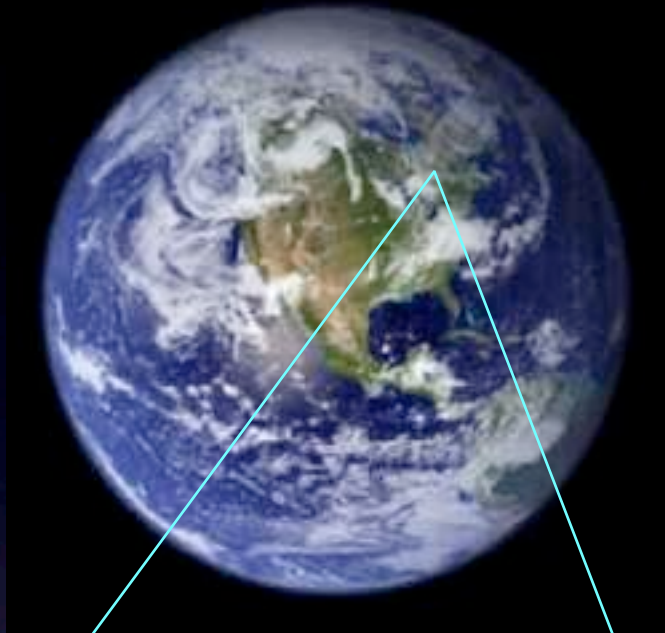
“How [have] we in astronomy come so far? ... By standing on the shoulders of military/ industrial giants. ... These larger scale efforts have been central to our success. ... Where military or industrial support did not exist and we had to go ahead on our own, progress has been much slower.”

Martin Harwit,
March 1999



The two main contractors that built the telescope had allegedly extensive experience building this kind of spacecraft - but not much is known publicly about these programs.

Large space based technologies rely on, and can enable multidisciplinary partnerships



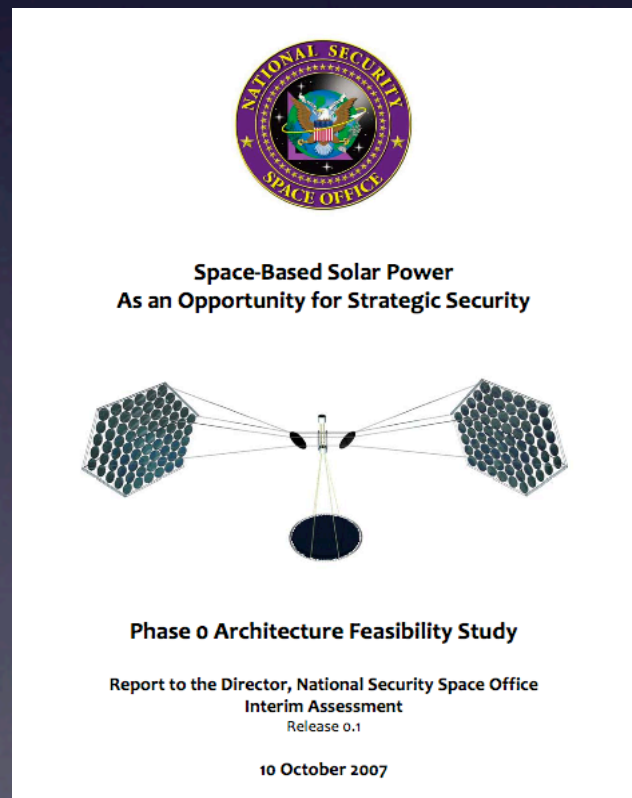
Geostationary orbit



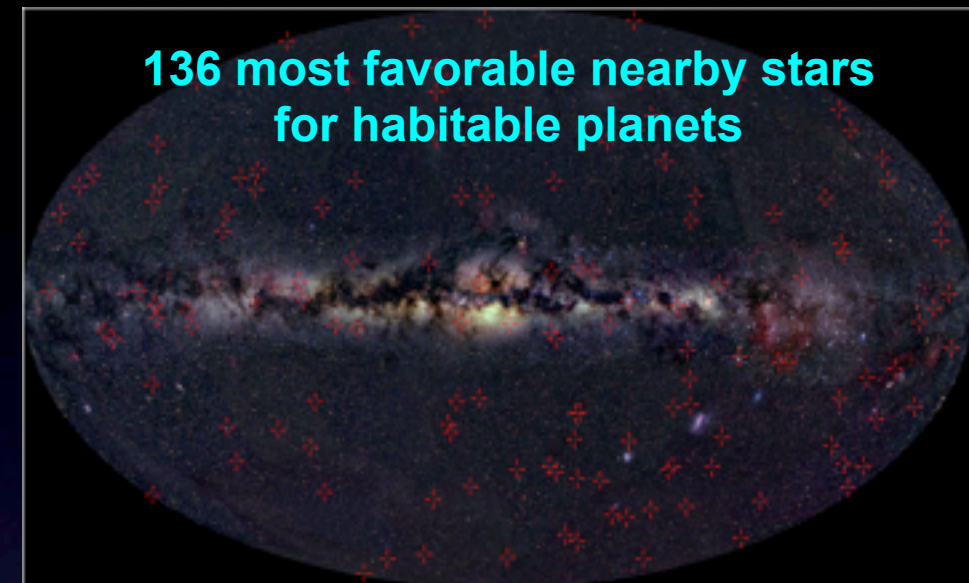
National & Environmental Security



Solar collectors in space



Energy



136 most favorable nearby stars for habitable planets

"I would like to see a reconnaissance of the planetary systems around the nearest 100 stars."

*Carl Sagan, 1994
(paraphrase)*

Are we alone?

Observable Drake Equation - *after Reid & Hawley*

$N_{L,T}$ is the number of life bearing planets at time T

$$N_{L,T} = N_{*,T} p_p n_e p_w p_a p_e p_s$$

Observable?

Number of stars @ T:

$$N_{*,T} = \text{SRF}(t) * \Psi(m) * \Lambda(m,t)$$



Prob. of planet system:

$$p_p = f(Z, m)$$



No. of terrestrial planets

$$n_e = n(0.1 m_e < m_p < 10 m_e)$$



Prob. of liquid water

$$p_w = f(n_b, \varepsilon, r_{\text{orbit}}, L^*)$$



Prob. of abiogenesis

$$p_a = f(?)$$



Prob. of evolution

$$p_e = f(t, Z, T)$$



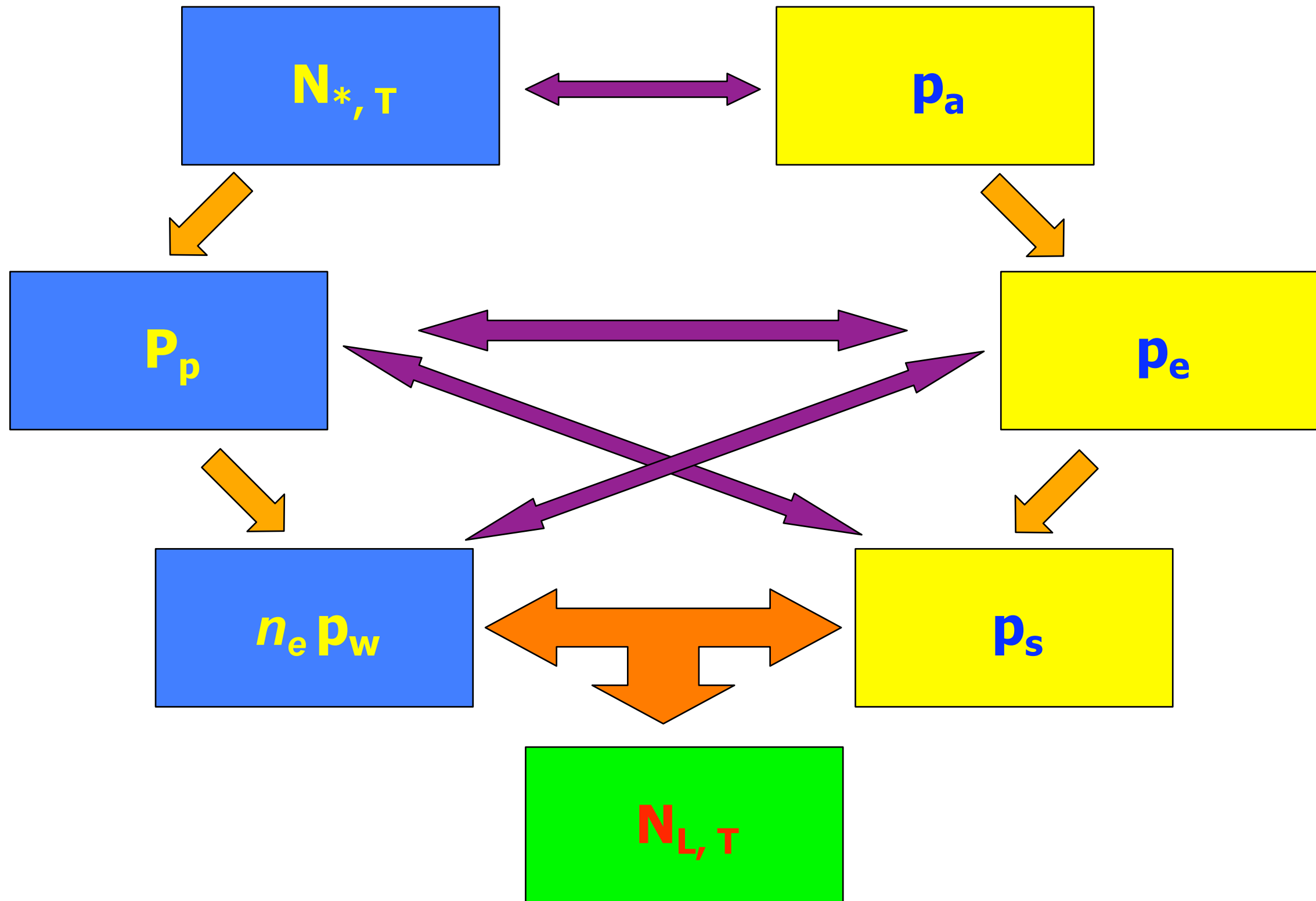
Prob. of survival

$$p_s = f(t, \text{environment})$$



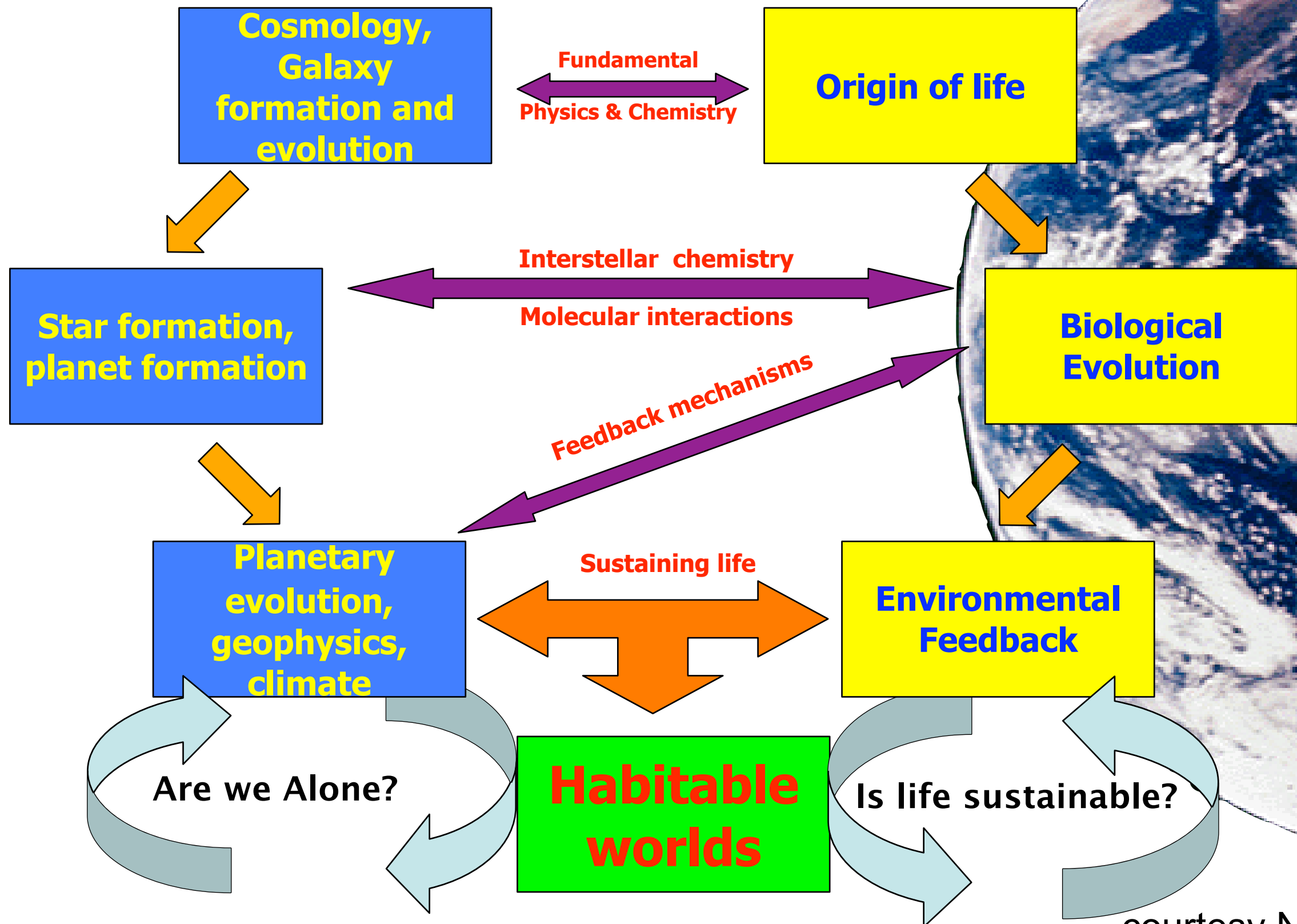
courtesy Neill Reid

Components of the Observable Drake Equation



courtesy Neill Reid

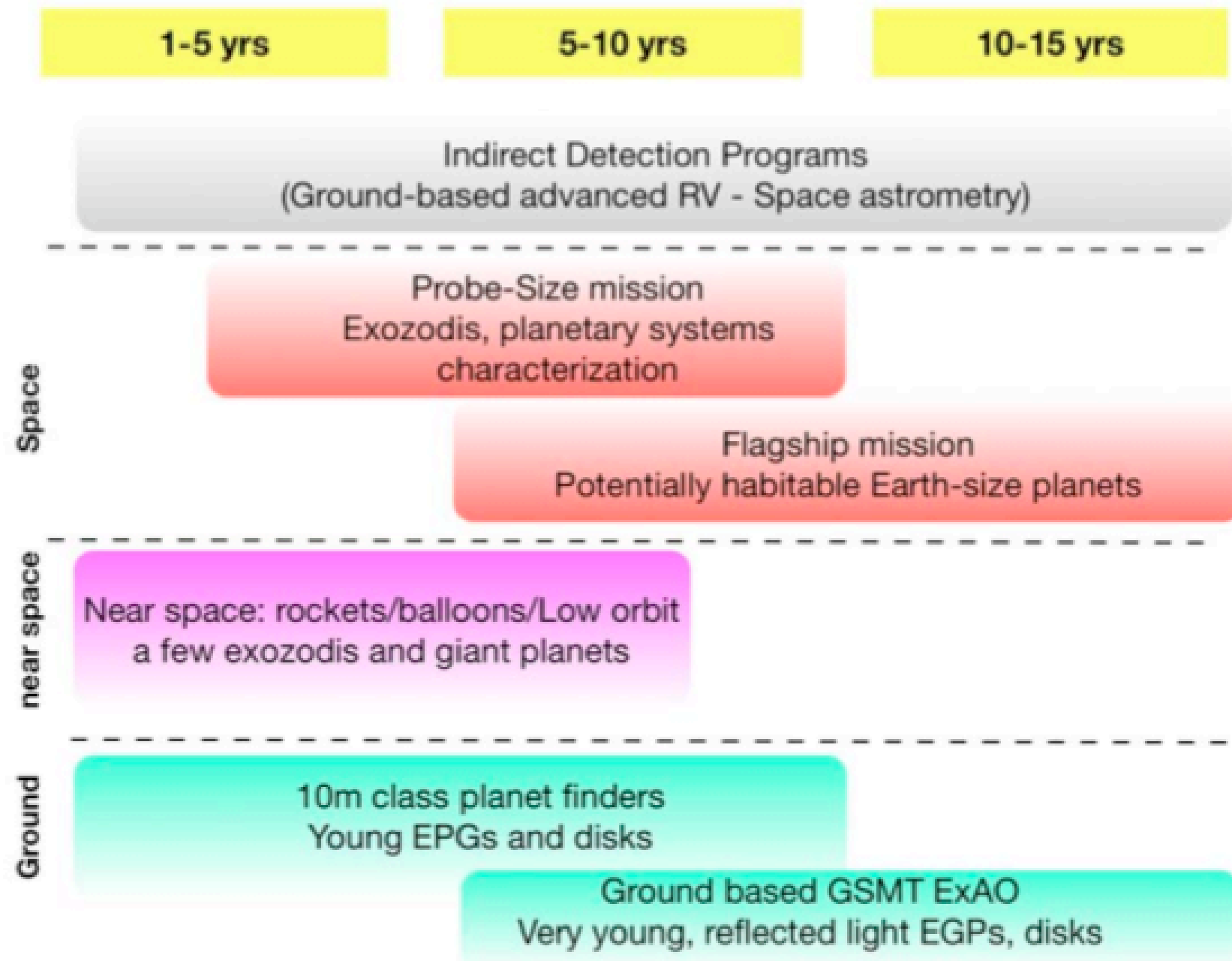
Components of the Observable Drake Equation



courtesy Neill Reid

Roadmap from “Exoplanet Community Report”

- Ed.by *P. R. Lawson, W.A. Traub and S. C. Unwin*



Note: boxes start with beginning of mission development

Figure 3-4. Range of relevant mission scales for direct imaging to enable a complete scientific and technological program. (R. Soummer, Space Telescope Science Institute, and M. Levine, JPL)

Roadmap from “Exoplanet Community Report”

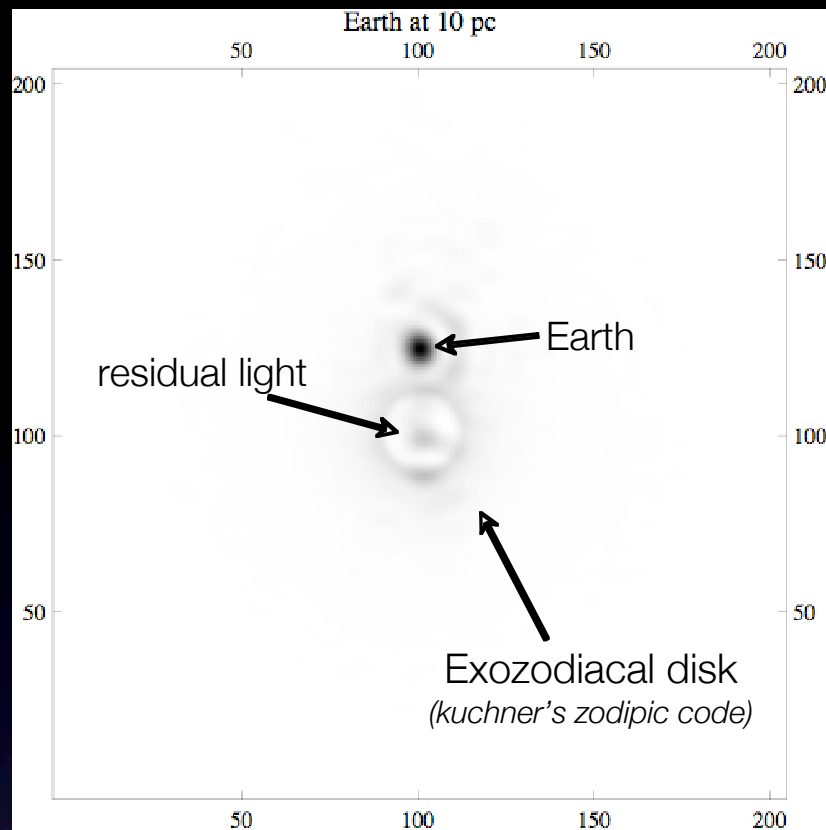
- Ed.by *P. R. Lawson, W.A. Traub and S. C. Unwin*



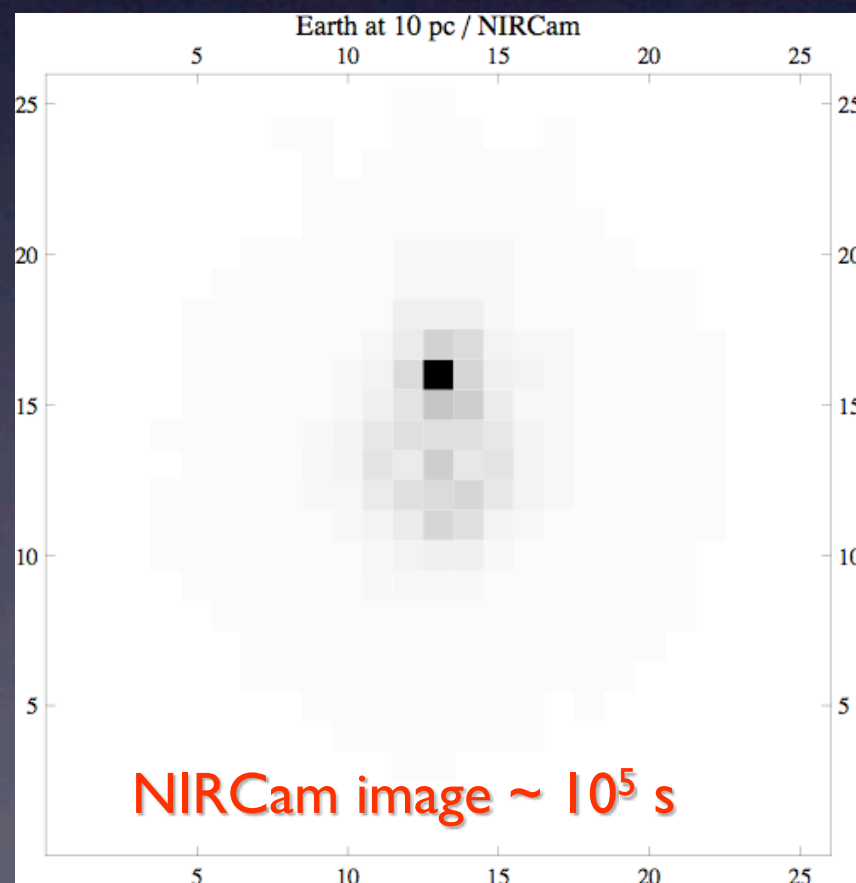
Figure 3-4. Range of relevant mission scales for direct imaging to enable a complete scientific and technological program. (R. Soummer, Space Telescope Science Institute, and M. Levine, JPL)

JWST with star-shade

- Tests a key technology
- Measures zodiacal light
- Can undertake spectroscopy of Jupiter's and “super-earths”
- Has the potential to deliver the first “pale blue dot”



Sun at 10pc with Earth $1e-10$ contrast, JWST's pupil and OPD



courtesy Remi Soummer, STScI

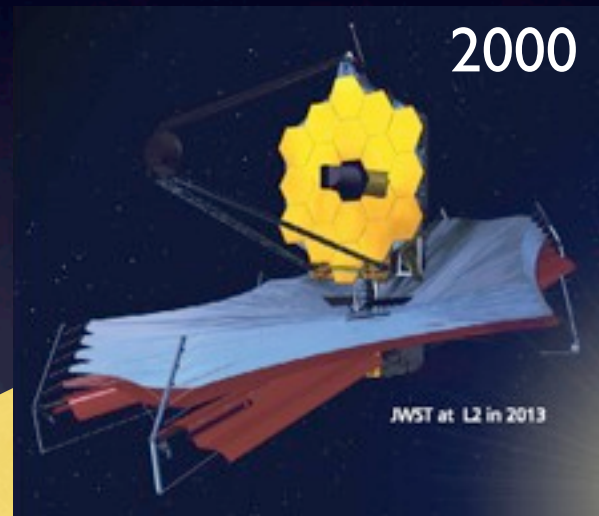
The Challenge

“Incrementalism is innovation’s worst enemy. We don’t want continuous improvement, we want radical change.”



1980

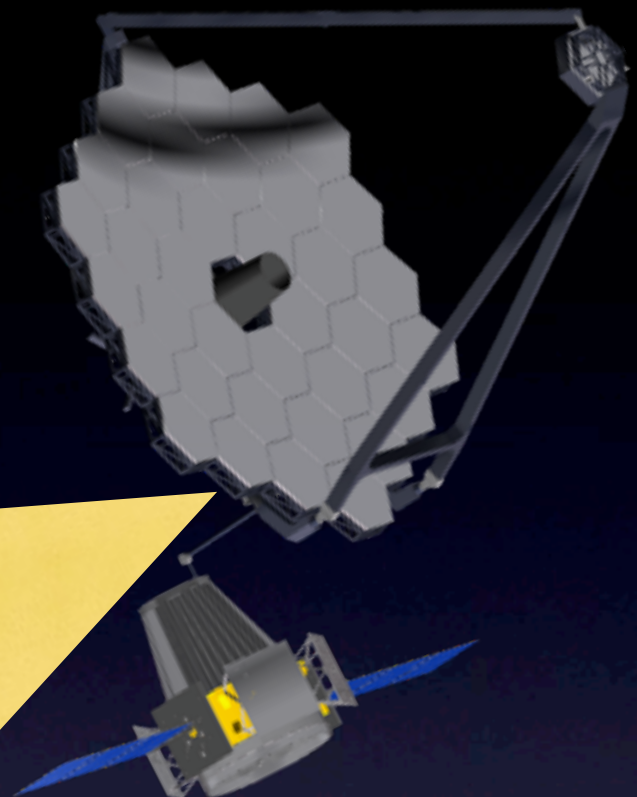
2.4m ~ \$4.5B (FY07)



2000

6.5m ~ \$4B (FY07)

2020



16m ~ \$4.5B (FY07)

focused investments,
and partnerships
should be able to deliver a
flagship class mission by 2020



within the next decade or so, we will have the capability to search for the signatures of life in our solar neighborhood.

we need to build a broad collation to support this endeavor, **then** our generation will be uniquely positioned “to relate causally the physical conditions during the Big Bang to the development of RNA and DNA¹.”

I. Riccardo Giacconi,

- Backup

NASA priorities in the next decade

The “Gap”



The Station



~\$100B

The Planet



Everything else...

NASA priorities in the next decade

The “Gap”



The Station



~\$100B

The Planet



To inspire the Nation

